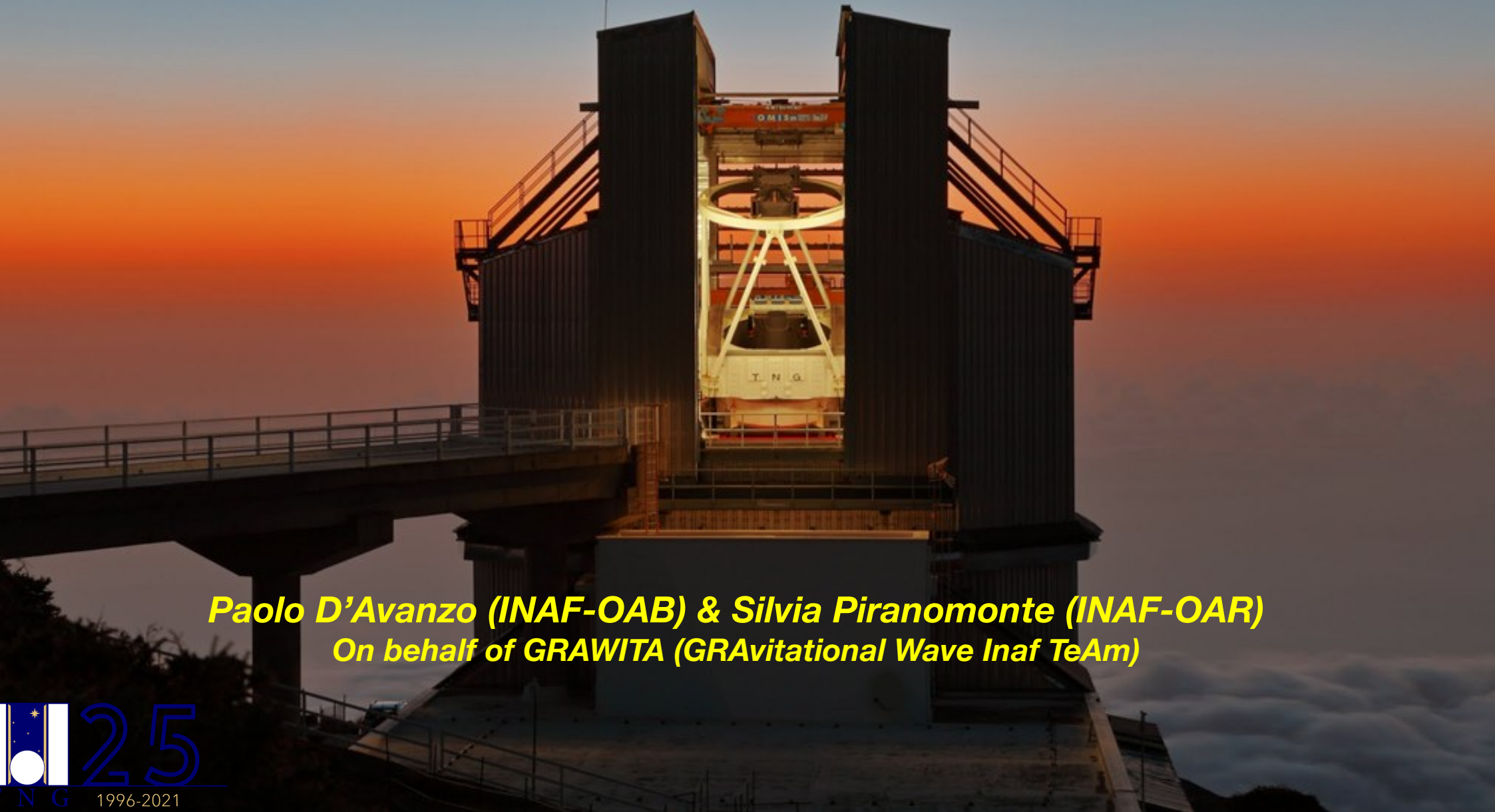
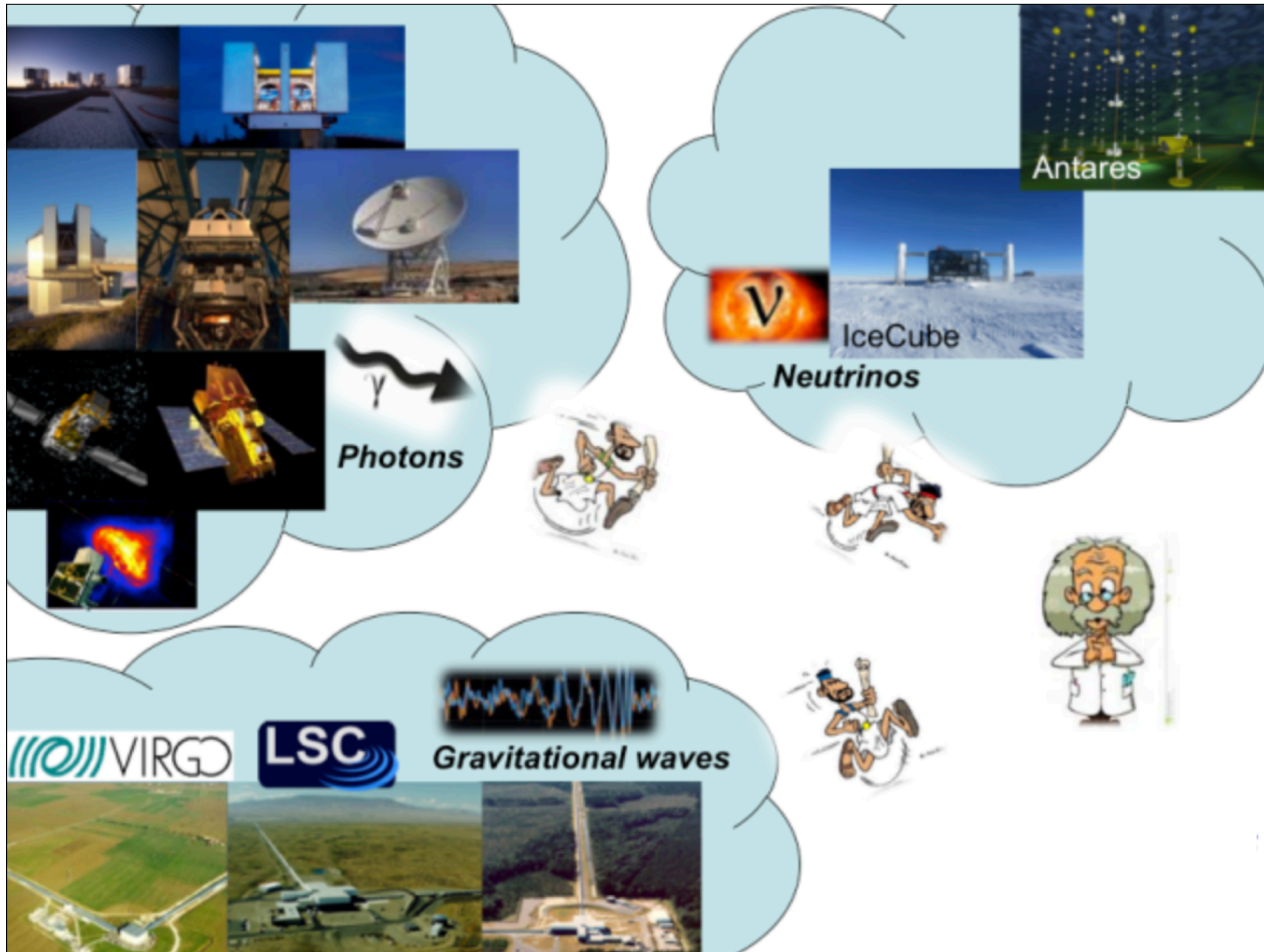
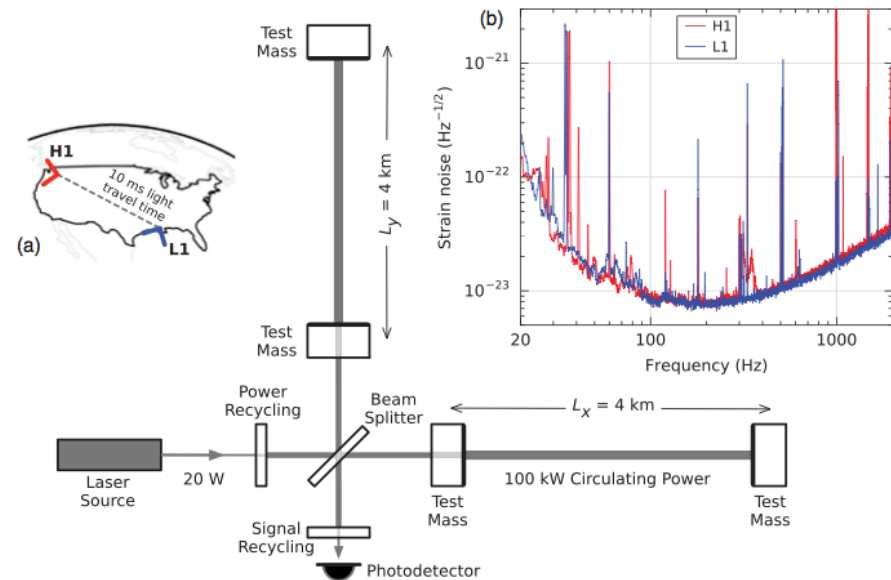
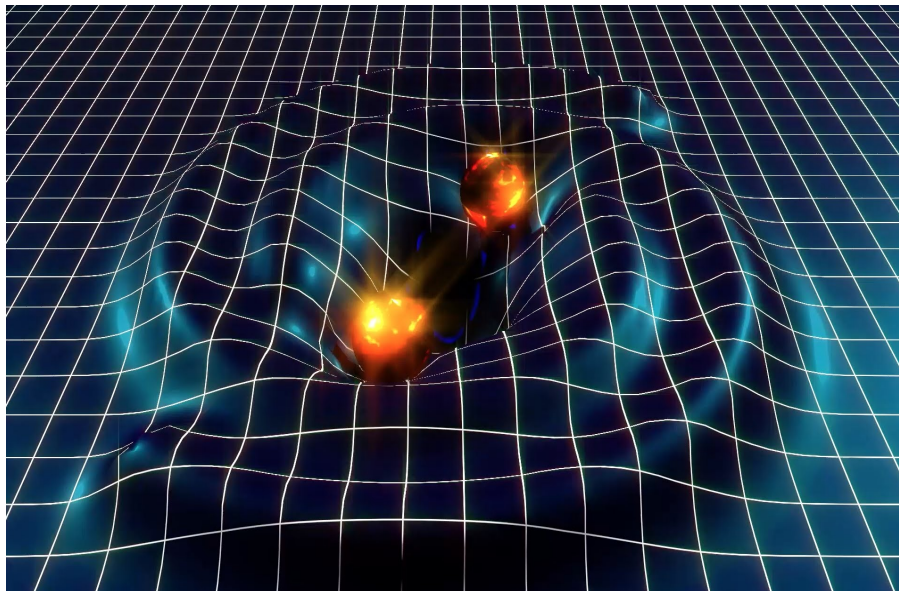


# Exploring the multi-messenger universe with TNG



***Paolo D'Avanzo (INAF-OAB) & Silvia Piranomonte (INAF-OAR)***  
***On behalf of GRAWITA (GRAvitational Wave Inaf TeAm)***







PRL 116, 061102 (2016)

Selected for a Viewpoint in Physics  
PHYSICAL REVIEW LETTERS

week ending  
12 FEBRUARY 2016



## Observation of Gravitational Waves from a Binary Black Hole Merger

B. P. Abbott *et al.*\*

(LIGO Scientific Collaboration and Virgo Collaboration)

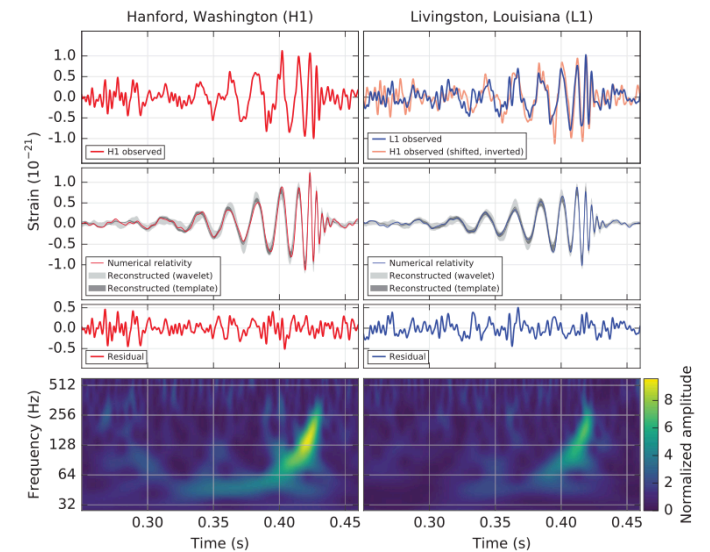
(Received 21 January 2016; published 11 February 2016)

On September 14, 2015 at 09:50:45 UTC the two detectors of the Laser Interferometer Gravitational-Wave Observatory simultaneously observed a transient gravitational-wave signal. The signal sweeps upwards in frequency from 35 to 250 Hz with a peak gravitational-wave strain of  $1.0 \times 10^{-21}$ . It matches the waveform predicted by general relativity for the inspiral and merger of a pair of black holes and the ringdown of the resulting single black hole. The signal was observed with a matched-filter signal-to-noise ratio of 24 and a false alarm rate estimated to be less than 1 event per 203 000 years, equivalent to a significance greater than  $5.1\sigma$ . The source lies at a luminosity distance of  $410^{+160}_{-180}$  Mpc corresponding to a redshift  $z = 0.09^{+0.03}_{-0.04}$ . In the source frame, the initial black hole masses are  $36^{+5}_{-4} M_{\odot}$  and  $29^{+4}_{-4} M_{\odot}$ , and the final black hole mass is  $62^{+4}_{-4} M_{\odot}$ , with  $3.0^{+0.5}_{-0.5} M_{\odot} c^2$  radiated in gravitational waves. All uncertainties define 90% credible intervals. These observations demonstrate the existence of binary stellar-mass black hole systems. This is the first direct detection of gravitational waves and the first observation of a binary black hole merger.

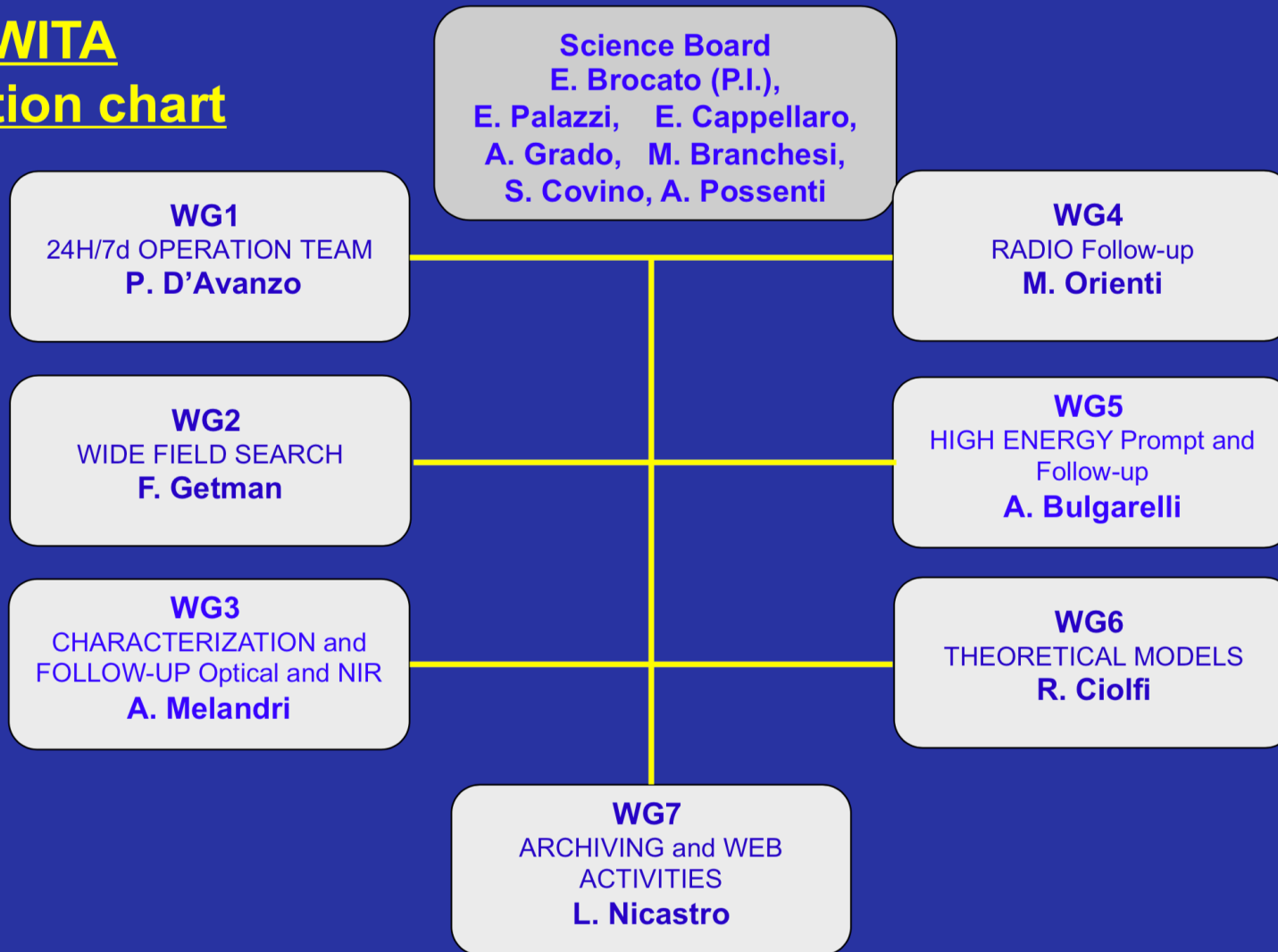
DOI: 10.1103/PhysRevLett.116.061102



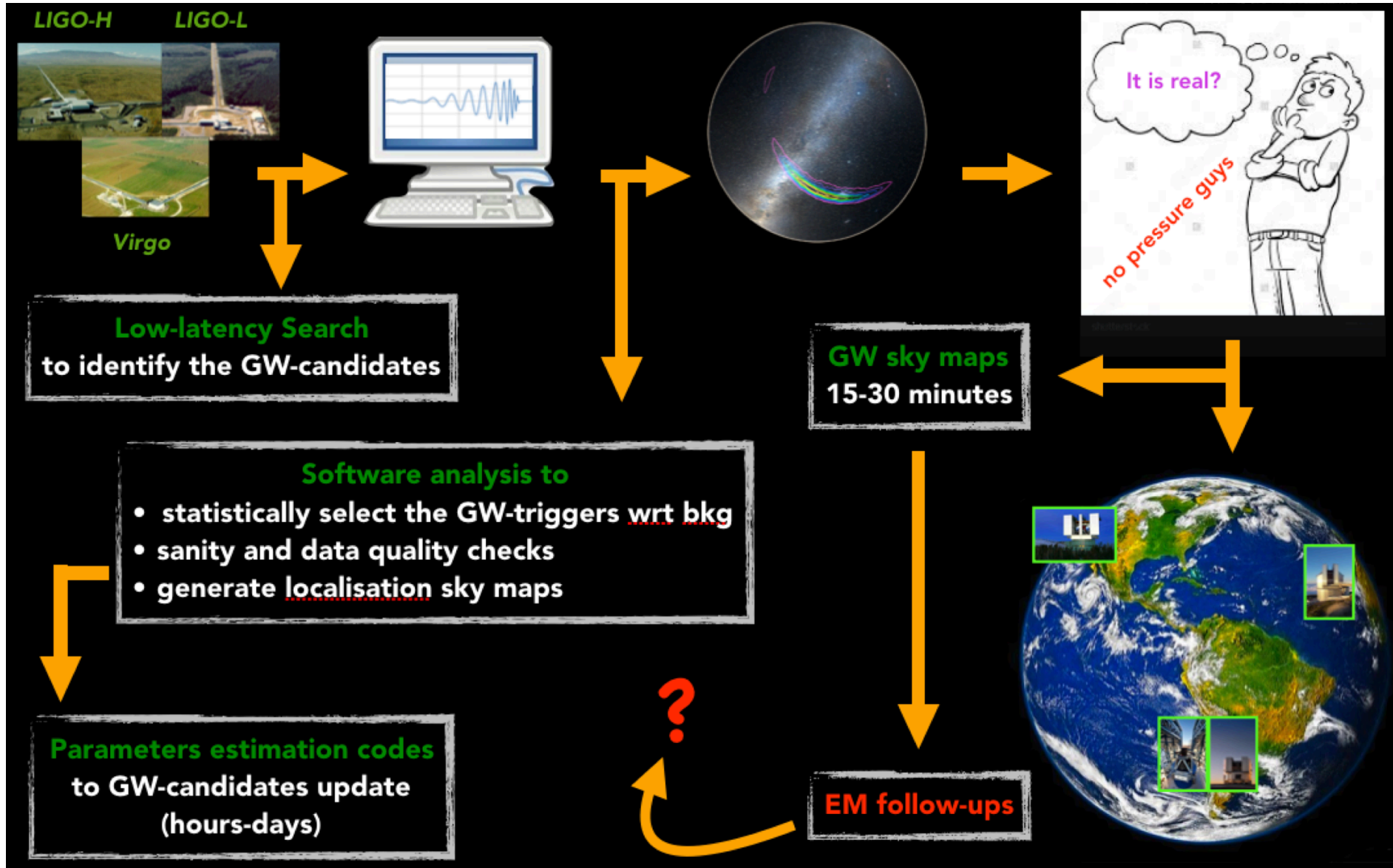
Primary black hole mass	$36^{+5}_{-4} M_{\odot}$
Secondary black hole mass	$29^{+4}_{-4} M_{\odot}$
Final black hole mass	$62^{+4}_{-4} M_{\odot}$
Final black hole spin	$0.67^{+0.05}_{-0.07}$
Luminosity distance	$410^{+160}_{-180}$ Mpc
Source redshift $z$	$0.09^{+0.03}_{-0.04}$

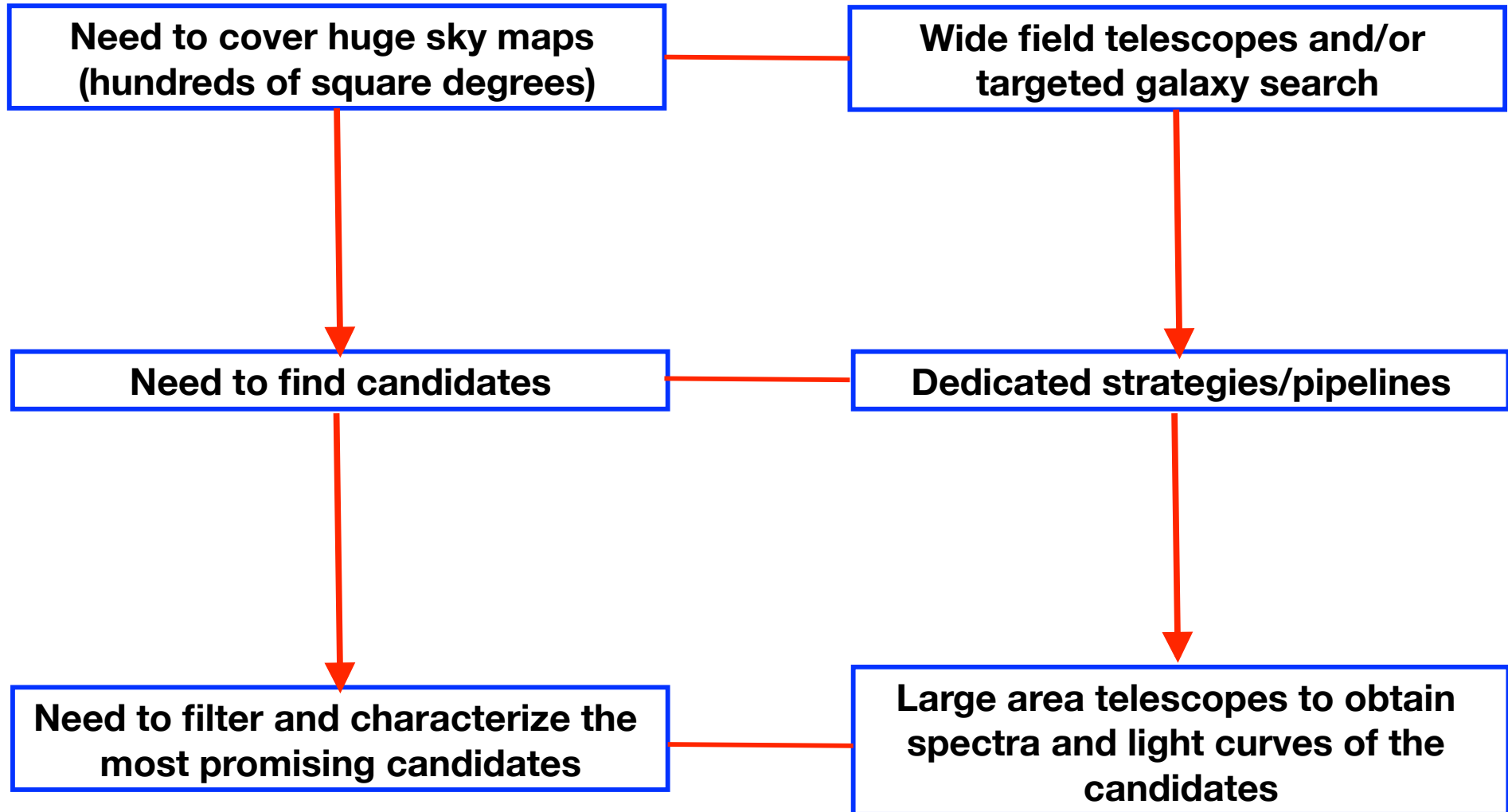


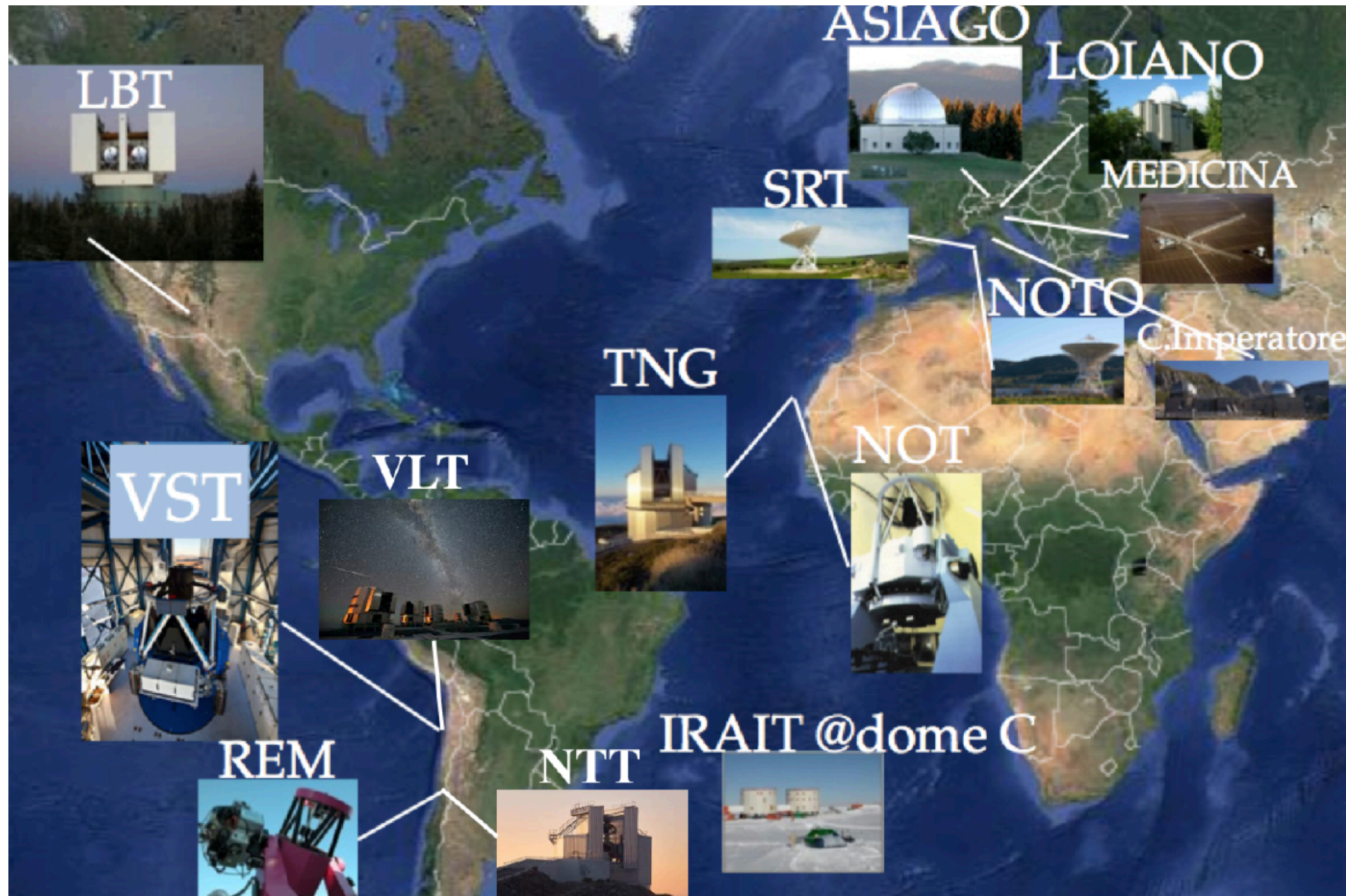
## GRAWITA Organization chart



~ 90 scientists from 21 Institutes (mainly INAF)



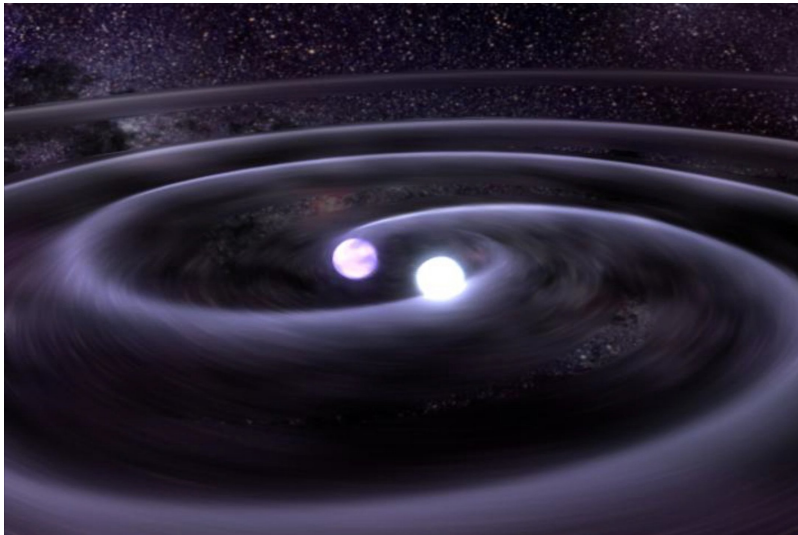




+ EVN & e-MERLIN

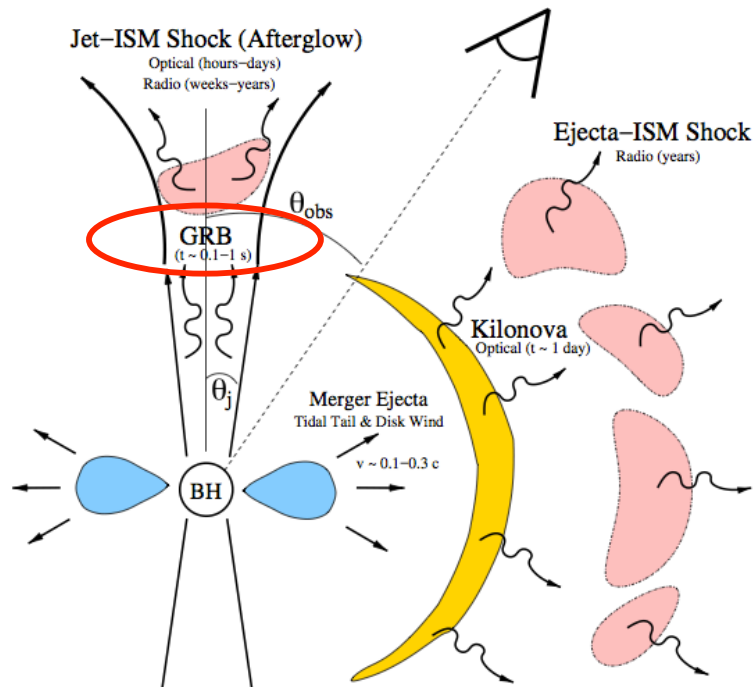
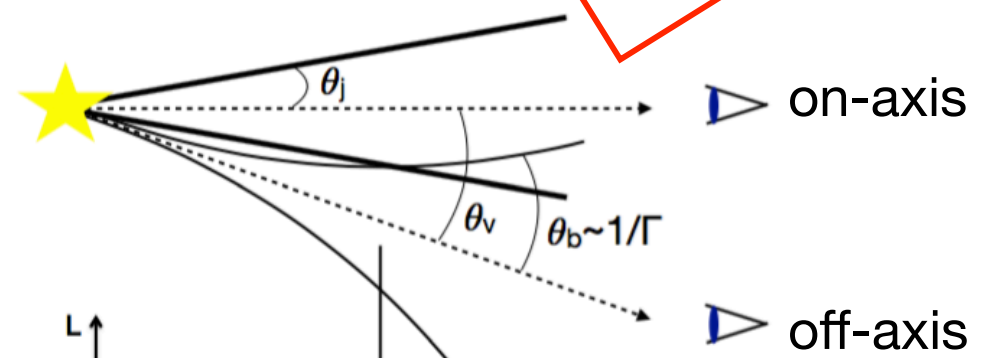
synergies with high-energy space observatories (Swift, AGILE, INTEGRAL, Fermi)



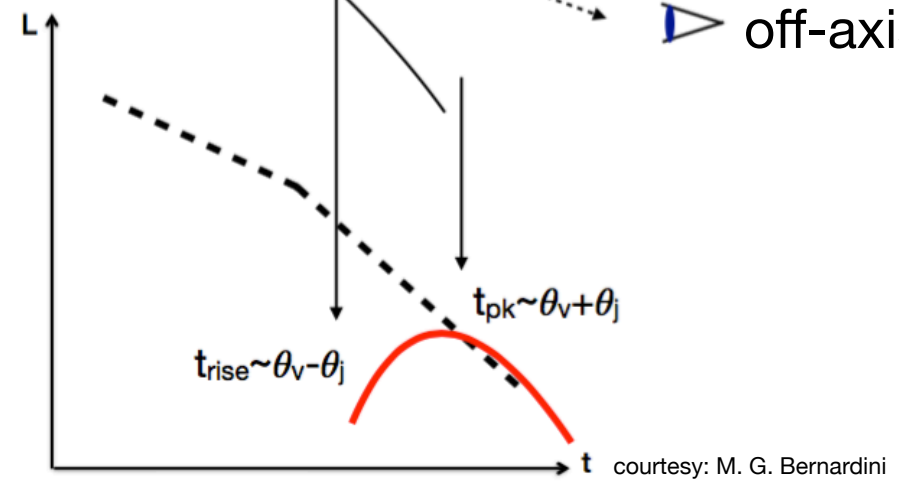


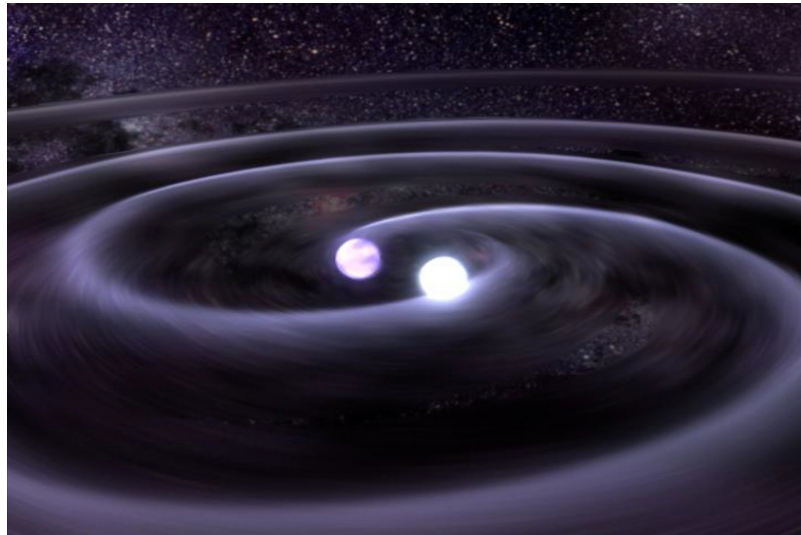
## Short GRB

Andrea Melandri's talk

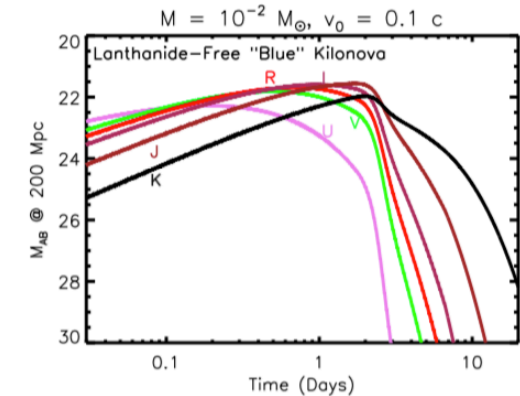
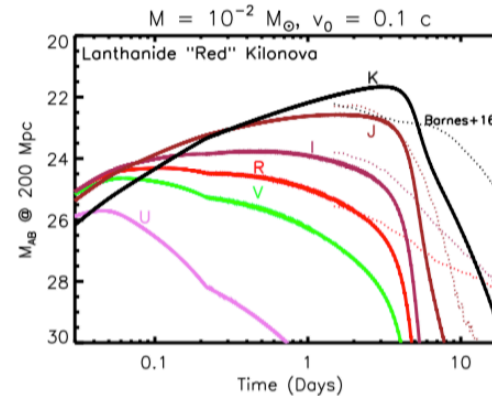


Metzger & Berger 2012

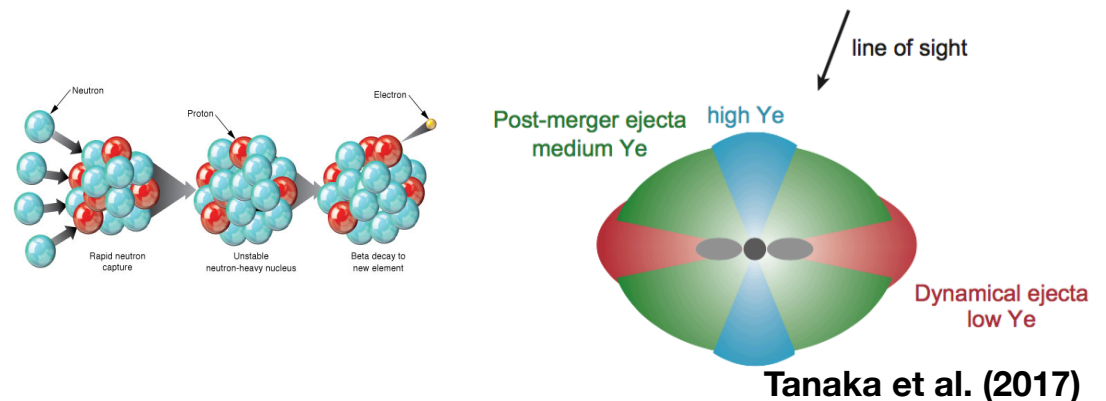
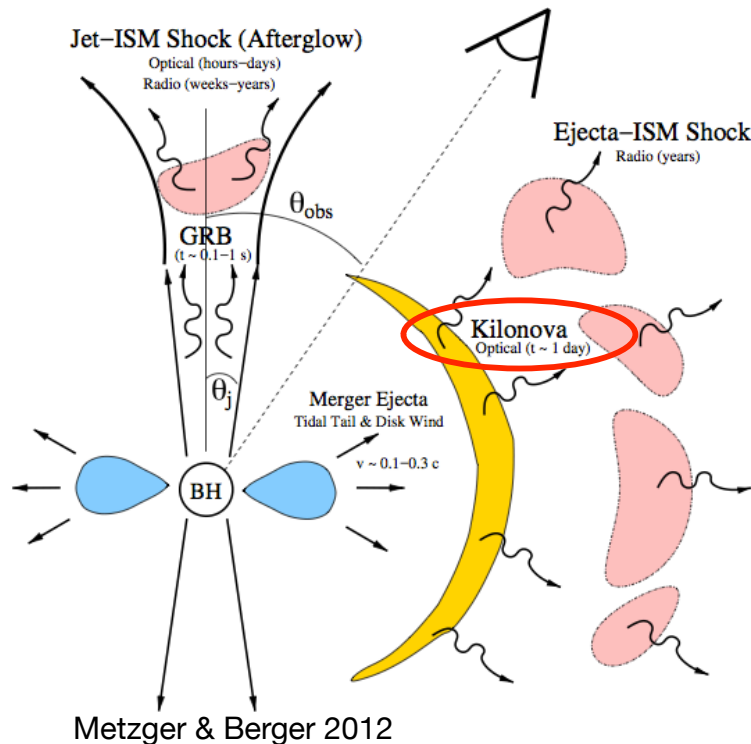




## Kilonova (aka macronova)



A key signature of an NS-NS/NS-BH binary merger is the production of a so-called **"kilonova"** (aka **"macronova"**) due to the decay of **heavy radioactive species** produced by the *r*-process and ejected during the merger that is expected to provide a source of heating and radiation (Li and Paczynski 1998; Rosswog, 2005; Metzger et al., 2010).







NA BEATA M ..... !!!



**Sept 2015 – Jan 2016: LVC O1 science run**  
**Nov 2016 – Aug 2017: LVC O2 science run**



Event	$m_1/M_\odot$	$m_2/M_\odot$	$M/M_\odot$	$\chi_{\text{eff}}$	$M_f/M_\odot$	$a_f$	$E_{\text{rad}}/(M_\odot c^2)$	$\ell_{\text{peak}}/(\text{erg s}^{-1})$	$d_L/\text{Mpc}$	$z$	$\Delta\Omega/\text{deg}^2$
GW150914	$35.6^{+4.8}_{-3.0}$	$30.6^{+3.0}_{-4.4}$	$28.6^{+1.6}_{-1.5}$	$-0.01^{+0.12}_{-0.13}$	$63.1^{+3.3}_{-3.0}$	$0.69^{+0.05}_{-0.04}$	$3.1^{+0.4}_{-0.4}$	$3.6^{+0.4}_{-0.4} \times 10^{56}$	$430^{+150}_{-170}$	$0.09^{+0.03}_{-0.03}$	180
GW151012	$23.3^{+14.0}_{-5.5}$	$13.6^{+4.1}_{-4.8}$	$15.2^{+2.0}_{-1.1}$	$0.04^{+0.28}_{-0.19}$	$35.7^{+9.9}_{-3.8}$	$0.67^{+0.13}_{-0.11}$	$1.5^{+0.5}_{-0.5}$	$3.2^{+0.8}_{-1.7} \times 10^{56}$	$1060^{+540}_{-480}$	$0.21^{+0.09}_{-0.09}$	1555
GW151226	$13.7^{+8.8}_{-3.2}$	$7.7^{+2.2}_{-2.6}$	$8.9^{+0.3}_{-0.3}$	$0.18^{+0.20}_{-0.12}$	$20.5^{+6.4}_{-1.5}$	$0.74^{+0.07}_{-0.05}$	$1.0^{+0.1}_{-0.2}$	$3.4^{+0.7}_{-1.7} \times 10^{56}$	$440^{+180}_{-190}$	$0.09^{+0.04}_{-0.04}$	1033
GW170104	$31.0^{+7.2}_{-5.6}$	$20.1^{+4.9}_{-4.5}$	$21.5^{+2.1}_{-1.7}$	$-0.04^{+0.17}_{-0.20}$	$49.1^{+5.2}_{-3.9}$	$0.66^{+0.08}_{-0.10}$	$2.2^{+0.5}_{-0.5}$	$3.3^{+0.6}_{-0.9} \times 10^{56}$	$960^{+430}_{-410}$	$0.19^{+0.07}_{-0.08}$	924
GW170608	$10.9^{+5.3}_{-1.7}$	$7.6^{+1.3}_{-2.1}$	$7.9^{+0.2}_{-0.2}$	$0.03^{+0.19}_{-0.07}$	$17.8^{+3.2}_{-0.7}$	$0.69^{+0.04}_{-0.04}$	$0.9^{+0.05}_{-0.1}$	$3.5^{+0.4}_{-1.3} \times 10^{56}$	$320^{+120}_{-110}$	$0.07^{+0.02}_{-0.02}$	396
GW170729	$50.6^{+16.6}_{-10.2}$	$34.3^{+9.1}_{-10.1}$	$35.7^{+6.5}_{-4.7}$	$0.36^{+0.21}_{-0.25}$	$80.3^{+14.6}_{-10.2}$	$0.81^{+0.07}_{-0.13}$	$4.8^{+1.7}_{-1.7}$	$4.2^{+0.9}_{-1.5} \times 10^{56}$	$2750^{+1350}_{-1320}$	$0.48^{+0.19}_{-0.20}$	1033
GW170809	$35.2^{+8.3}_{-6.0}$	$23.8^{+5.2}_{-5.1}$	$25.0^{+2.1}_{-1.6}$	$0.07^{+0.16}_{-0.16}$	$56.4^{+5.2}_{-3.7}$	$0.70^{+0.08}_{-0.09}$	$2.7^{+0.6}_{-0.6}$	$3.5^{+0.6}_{-0.9} \times 10^{56}$	$990^{+320}_{-380}$	$0.20^{+0.05}_{-0.07}$	340
GW170814	$30.7^{+5.7}_{-3.0}$	$25.3^{+2.9}_{-4.1}$	$24.2^{+1.4}_{-1.1}$	$0.07^{+0.12}_{-0.11}$	$53.4^{+3.2}_{-2.4}$	$0.72^{+0.07}_{-0.05}$	$2.7^{+0.4}_{-0.3}$	$3.7^{+0.4}_{-0.5} \times 10^{56}$	$580^{+160}_{-210}$	$0.12^{+0.03}_{-0.04}$	87
GW170817	$1.46^{+0.12}_{-0.10}$	$1.27^{+0.09}_{-0.09}$	$1.186^{+0.001}_{-0.001}$	$0.00^{+0.02}_{-0.01}$	$\leq 2.8$	$\leq 0.89$	$\geq 0.04$	$\geq 0.1 \times 10^{56}$	$40^{+10}_{-10}$	$0.01^{+0.00}_{-0.00}$	16
GW170818	$35.5^{+7.5}_{-4.7}$	$26.8^{+4.3}_{-5.2}$	$26.7^{+2.1}_{-1.7}$	$-0.09^{+0.18}_{-0.21}$	$59.8^{+4.8}_{-3.8}$	$0.67^{+0.07}_{-0.08}$	$2.7^{+0.5}_{-0.5}$	$3.4^{+0.5}_{-0.7} \times 10^{56}$	$1020^{+430}_{-360}$	$0.20^{+0.07}_{-0.07}$	39
GW170823	$39.6^{+10.0}_{-6.6}$	$29.4^{+6.3}_{-7.1}$	$29.3^{+4.2}_{-3.2}$	$0.08^{+0.20}_{-0.22}$	$65.6^{+9.4}_{-6.6}$	$0.71^{+0.08}_{-0.10}$	$3.3^{+0.9}_{-0.8}$	$3.6^{+0.6}_{-0.9} \times 10^{56}$	$1850^{+840}_{-840}$	$0.34^{+0.13}_{-0.14}$	1651





**Sept 2015 – Jan 2016: LVC O1 science run**  
**Nov 2016 – Aug 2017: LVC O2 science run**



Event	$m_1/M_\odot$	$m_2/M_\odot$	$M/M_\odot$	$\chi_{\text{eff}}$	$M_f/M_\odot$	$a_f$	$E_{\text{rad}}/(M_\odot c^2)$	$\ell_{\text{peak}}/(\text{erg s}^{-1})$	$d_L/\text{Mpc}$	$z$	$\Delta\Omega/\text{deg}^2$
<a href="#">GW150914</a>	$35.6^{+4.8}_{-3.0}$	$30.6^{+3.0}_{-4.4}$	$28.6^{+1.6}_{-1.5}$	$-0.01^{+0.12}_{-0.13}$	$63.1^{+3.3}_{-3.0}$	$0.69^{+0.05}_{-0.04}$	$3.1^{+0.4}_{-0.4}$	$3.6^{+0.4}_{-0.4} \times 10^{56}$	$430^{+150}_{-170}$	$0.09^{+0.03}_{-0.03}$	180
<a href="#">GW151012</a>	$23.3^{+14.0}_{-5.5}$	$13.6^{+4.1}_{-4.8}$	$15.2^{+2.0}_{-1.1}$	$0.04^{+0.28}_{-0.19}$	$35.7^{+9.9}_{-3.8}$	$0.67^{+0.13}_{-0.11}$	$1.5^{+0.5}_{-0.5}$	$3.2^{+0.8}_{-1.7} \times 10^{56}$	$1060^{+540}_{-480}$	$0.21^{+0.09}_{-0.09}$	1555
<a href="#">GW151226</a>	$13.7^{+8.8}_{-3.2}$	$7.7^{+2.2}_{-2.6}$	$8.9^{+0.3}_{-0.3}$	$0.18^{+0.20}_{-0.12}$	$20.5^{+6.4}_{-1.5}$	$0.74^{+0.07}_{-0.05}$	$1.0^{+0.1}_{-0.2}$	$3.4^{+0.7}_{-1.7} \times 10^{56}$	$440^{+180}_{-190}$	$0.09^{+0.04}_{-0.04}$	1033
<a href="#">GW170104</a>	$31.0^{+7.2}_{-5.6}$	$20.1^{+4.9}_{-4.5}$	$21.5^{+2.1}_{-1.7}$	$-0.04^{+0.17}_{-0.20}$	$49.1^{+5.2}_{-3.9}$	$0.66^{+0.08}_{-0.10}$	$2.2^{+0.5}_{-0.5}$	$3.3^{+0.6}_{-0.9} \times 10^{56}$	$960^{+430}_{-410}$	$0.19^{+0.07}_{-0.08}$	924
<a href="#">GW170608</a>	$10.9^{+5.3}_{-1.7}$	$7.6^{+1.3}_{-2.1}$	$7.9^{+0.2}_{-0.2}$	$0.03^{+0.19}_{-0.07}$	$17.8^{+3.2}_{-0.7}$	$0.69^{+0.04}_{-0.04}$	$0.9^{+0.05}_{-0.1}$	$3.5^{+0.4}_{-1.3} \times 10^{56}$	$320^{+120}_{-110}$	$0.07^{+0.02}_{-0.02}$	396
<a href="#">GW170729</a>	$50.6^{+16.6}_{-10.2}$	$34.3^{+9.1}_{-10.1}$	$35.7^{+6.5}_{-4.7}$	$0.36^{+0.21}_{-0.25}$	$80.3^{+14.6}_{-10.2}$	$0.81^{+0.07}_{-0.13}$	$4.8^{+1.7}_{-1.7}$	$4.2^{+0.9}_{-1.5} \times 10^{56}$	$2750^{+1350}_{-1320}$	$0.48^{+0.19}_{-0.20}$	1033
<a href="#">GW170809</a>	$35.2^{+8.3}_{-6.0}$	$23.8^{+5.2}_{-5.1}$	$25.0^{+2.1}_{-1.6}$	$0.07^{+0.16}_{-0.16}$	$56.4^{+5.2}_{-3.7}$	$0.70^{+0.08}_{-0.09}$	$2.7^{+0.6}_{-0.6}$	$3.5^{+0.6}_{-0.9} \times 10^{56}$	$990^{+320}_{-380}$	$0.20^{+0.05}_{-0.07}$	340
<a href="#">GW170814</a>	$30.7^{+5.7}_{-3.0}$	$25.3^{+2.9}_{-4.1}$	$24.2^{+1.4}_{-1.1}$	$0.07^{+0.12}_{-0.11}$	$53.4^{+3.2}_{-2.4}$	$0.72^{+0.07}_{-0.05}$	$2.7^{+0.4}_{-0.3}$	$3.7^{+0.4}_{-0.5} \times 10^{56}$	$580^{+160}_{-210}$	$0.12^{+0.03}_{-0.04}$	87
<a href="#">GW170817</a>	$1.46^{+0.12}_{-0.10}$	$1.27^{+0.09}_{-0.09}$	$1.186^{+0.001}_{-0.001}$	$0.00^{+0.02}_{-0.01}$	$\leq 2.8$	$\leq 0.89$	$\geq 0.04$	$\geq 0.1 \times 10^{56}$	$40^{+10}_{-10}$	$0.01^{+0.00}_{-0.00}$	16
<a href="#">GW170818</a>	$35.5^{+7.5}_{-4.7}$	$26.8^{+4.3}_{-5.2}$	$26.7^{+2.1}_{-1.7}$	$-0.09^{+0.18}_{-0.21}$	$59.8^{+4.8}_{-3.8}$	$0.67^{+0.07}_{-0.08}$	$2.7^{+0.5}_{-0.5}$	$3.4^{+0.5}_{-0.7} \times 10^{56}$	$1020^{+430}_{-360}$	$0.20^{+0.07}_{-0.07}$	39
<a href="#">GW170823</a>	$39.6^{+10.0}_{-6.6}$	$29.4^{+6.3}_{-7.1}$	$29.3^{+4.2}_{-3.2}$	$0.08^{+0.20}_{-0.22}$	$65.6^{+9.4}_{-6.6}$	$0.71^{+0.08}_{-0.10}$	$3.3^{+0.9}_{-0.8}$	$3.6^{+0.6}_{-0.9} \times 10^{56}$	$1850^{+840}_{-840}$	$0.34^{+0.13}_{-0.14}$	1651



**GW 150914:** 630 deg<sup>2</sup> skymap  
 90 deg<sup>2</sup> observed with VST  
 (6 epochs between Sep 17 and Nov 18)

**Brocato+16; Abbott+16**

**GW 151226:** 1240 deg<sup>2</sup> skymap  
 72 deg<sup>2</sup> observed with VST  
 (6 epochs between Dec 27 and Feb 10)

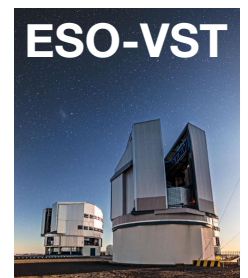
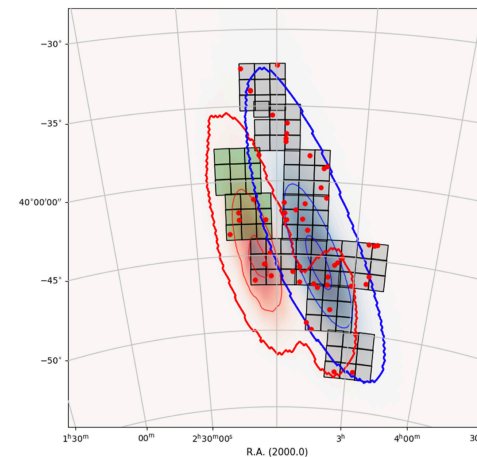
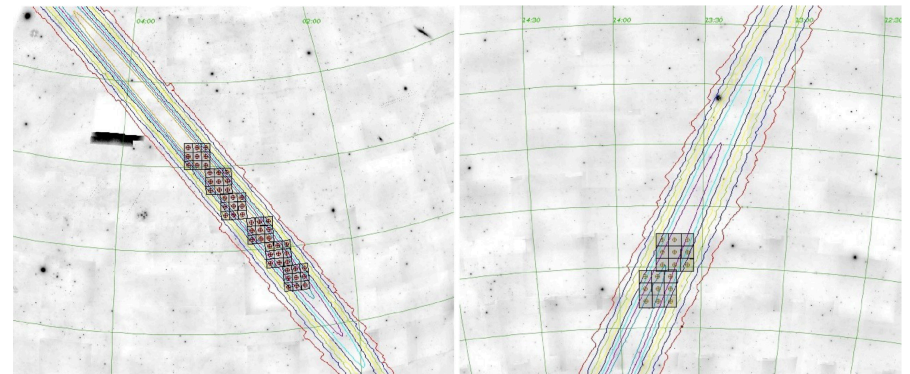
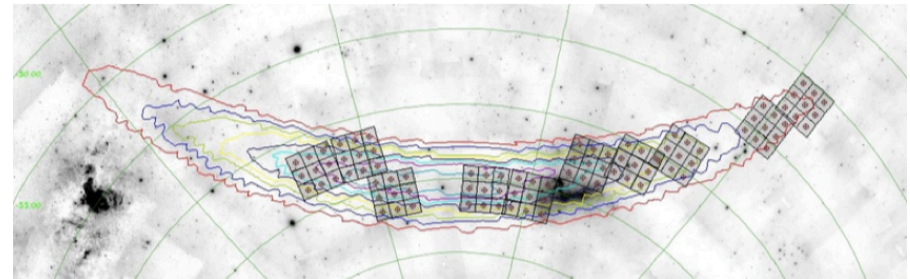
**Brocato+16**

**GW 170814:** 87 deg<sup>2</sup> skymap  
**59%** observed with VST  
 (6 epochs between Aug 14 and Sep 28)

**Grado+20**

Huge observational effort, mainly with wide-field facilities

**Several transients discovered**  
**No EM counterpart found**





2015GCN.18536....1D2015

LIGO/Virgo G194575: **INAF-TNG** follow-up of iPTF15dkm and iPTF15dlj

D'Avanzo, P.; Melandri, A.; Piranomonte, S. and 28 more

2015GCN.18488....1P2015

LIGO/Virgo G194575: **INAF-TNG** spectra of LSQ15bjb

Piranomonte, S.; D'Avanzo, P.; Melandri, A. and 23 more

2015GCN.18476....1D2015

LIGO/Virgo G194575: **INAF-TNG** follow-up of LSQ15bjb

D'Avanzo, P.; Melandri, A.; Piranomonte, S. and 19 more

2015GCN.18775....1D2015

LIGO/Virgo G211117: **INAF TNG** follow-up of MASTER OT J020906.21+013800.1.

D'Avanzo, P.; Melandri, A.; Piranomonte, S. and 28 more

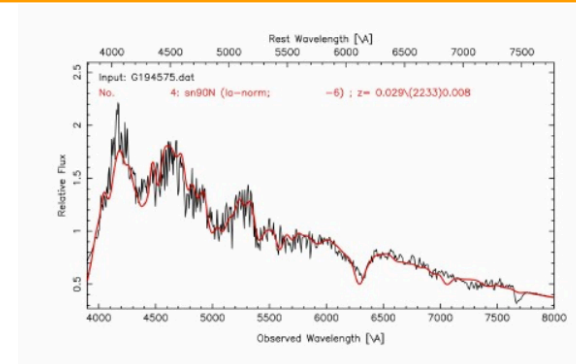


Figure 1: TNG spectrum of LSQ15bjb of 26 October 2015 (black) compared with that of SN 1990N (red), six days before *B*-band maximum light.

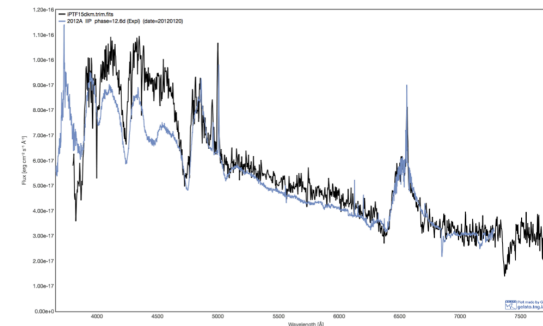


Figure 2: TNG spectrum of iPTF15dkm of 29 October 2015 (black) and best match with SN 2012A (blue), using GELATO (Harutyunyan et al. 2008, A&A, 488, 383).

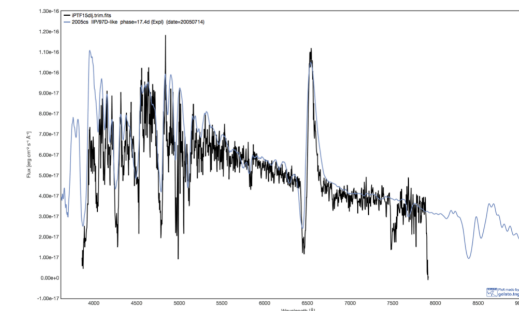
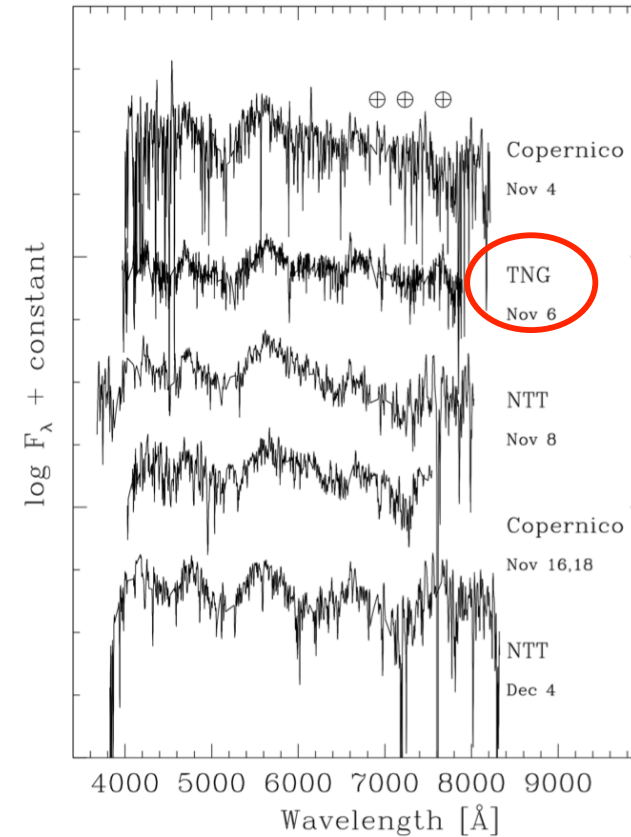
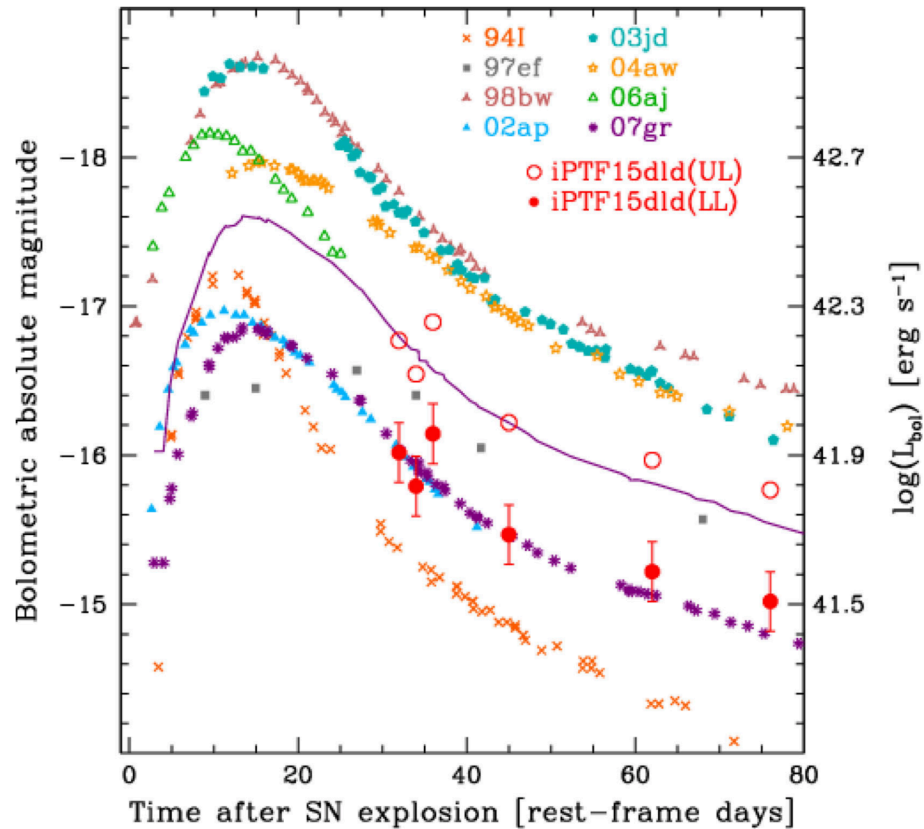


Figure 3: TNG spectrum of iPTF15dlj of 29 October 2015 (black) and best match with SN 2005cs (blue), using GELATO (Harutyunyan et al. 2008, A&A, 488, 383).





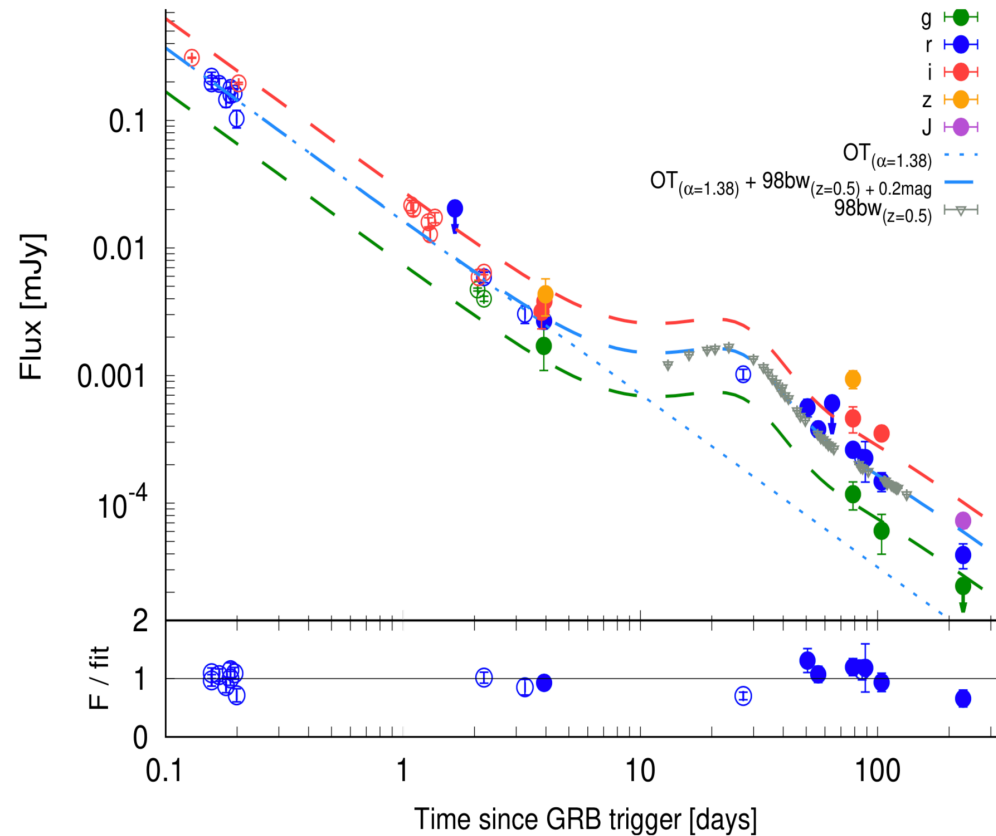
**iPTF15dld:**  
a low-luminosity, broad-line  
Ic supernova  
**Pian+17**





**ATLAS17aeu:**  
 transient discovered during the search for the e.m. counterpart of GW 170104 (BBH). Likely associated with the long GRB 170105 (and associated SN) at  $z_{\text{ph}} \sim 0.5$

**Melandri+19**





**Sept 2015 – Jan 2016: LVC O1 science run**  
**Nov 2016 – Aug 2017: LVC O2 science run**



Event	$m_1/M_\odot$	$m_2/M_\odot$	$M/M_\odot$	$\chi_{\text{eff}}$	$M_f/M_\odot$	$a_f$	$E_{\text{rad}}/(M_\odot c^2)$	$\ell_{\text{peak}}/(\text{erg s}^{-1})$	$d_L/\text{Mpc}$	$z$	$\Delta\Omega/\text{deg}^2$
<a href="#">GW150914</a>	$35.6^{+4.8}_{-3.0}$	$30.6^{+3.0}_{-4.4}$	$28.6^{+1.6}_{-1.5}$	$-0.01^{+0.12}_{-0.13}$	$63.1^{+3.3}_{-3.0}$	$0.69^{+0.05}_{-0.04}$	$3.1^{+0.4}_{-0.4}$	$3.6^{+0.4}_{-0.4} \times 10^{56}$	$430^{+150}_{-170}$	$0.09^{+0.03}_{-0.03}$	180
<a href="#">GW151012</a>	$23.3^{+14.0}_{-5.5}$	$13.6^{+4.1}_{-4.8}$	$15.2^{+2.0}_{-1.1}$	$0.04^{+0.28}_{-0.19}$	$35.7^{+9.9}_{-3.8}$	$0.67^{+0.13}_{-0.11}$	$1.5^{+0.5}_{-0.5}$	$3.2^{+0.8}_{-1.7} \times 10^{56}$	$1060^{+540}_{-480}$	$0.21^{+0.09}_{-0.09}$	1555
<a href="#">GW151226</a>	$13.7^{+8.8}_{-3.2}$	$7.7^{+2.2}_{-2.6}$	$8.9^{+0.3}_{-0.3}$	$0.18^{+0.20}_{-0.12}$	$20.5^{+6.4}_{-1.5}$	$0.74^{+0.07}_{-0.05}$	$1.0^{+0.1}_{-0.2}$	$3.4^{+0.7}_{-1.7} \times 10^{56}$	$440^{+180}_{-190}$	$0.09^{+0.04}_{-0.04}$	1033
<a href="#">GW170104</a>	$31.0^{+7.2}_{-5.6}$	$20.1^{+4.9}_{-4.5}$	$21.5^{+2.1}_{-1.7}$	$-0.04^{+0.17}_{-0.20}$	$49.1^{+5.2}_{-3.9}$	$0.66^{+0.08}_{-0.10}$	$2.2^{+0.5}_{-0.5}$	$3.3^{+0.6}_{-0.9} \times 10^{56}$	$960^{+430}_{-410}$	$0.19^{+0.07}_{-0.08}$	924
<a href="#">GW170608</a>	$10.9^{+5.3}_{-1.7}$	$7.6^{+1.3}_{-2.1}$	$7.9^{+0.2}_{-0.2}$	$0.03^{+0.19}_{-0.07}$	$17.8^{+3.2}_{-0.7}$	$0.69^{+0.04}_{-0.04}$	$0.9^{+0.05}_{-0.1}$	$3.5^{+0.4}_{-1.3} \times 10^{56}$	$320^{+120}_{-110}$	$0.07^{+0.02}_{-0.02}$	396
<a href="#">GW170729</a>	$50.6^{+16.6}_{-10.2}$	$34.3^{+9.1}_{-10.1}$	$35.7^{+6.5}_{-4.7}$	$0.36^{+0.21}_{-0.25}$	$80.3^{+14.6}_{-10.2}$	$0.81^{+0.07}_{-0.13}$	$4.8^{+1.7}_{-1.7}$	$4.2^{+0.9}_{-1.5} \times 10^{56}$	$2750^{+1350}_{-1320}$	$0.48^{+0.19}_{-0.20}$	1033
<a href="#">GW170809</a>	$35.2^{+8.3}_{-6.0}$	$23.8^{+5.2}_{-5.1}$	$25.0^{+2.1}_{-1.6}$	$0.07^{+0.16}_{-0.16}$	$56.4^{+5.2}_{-3.7}$	$0.70^{+0.08}_{-0.09}$	$2.7^{+0.6}_{-0.6}$	$3.5^{+0.6}_{-0.9} \times 10^{56}$	$990^{+320}_{-380}$	$0.20^{+0.05}_{-0.07}$	340
<a href="#">GW170814</a>	$30.7^{+5.7}_{-3.0}$	$25.3^{+2.9}_{-4.1}$	$24.2^{+1.4}_{-1.1}$	$0.07^{+0.12}_{-0.11}$	$53.4^{+3.2}_{-2.4}$	$0.72^{+0.07}_{-0.05}$	$2.7^{+0.4}_{-0.3}$	$3.7^{+0.4}_{-0.5} \times 10^{56}$	$580^{+160}_{-210}$	$0.12^{+0.03}_{-0.04}$	87
<a href="#">GW170817</a>	$1.46^{+0.12}_{-0.10}$	$1.27^{+0.09}_{-0.09}$	$1.186^{+0.001}_{-0.001}$	$0.00^{+0.02}_{-0.01}$	$\leq 2.8$	$\leq 0.89$	$\geq 0.04$	$\geq 0.1 \times 10^{56}$	$40^{+10}_{-10}$	$0.01^{+0.00}_{-0.00}$	16
<a href="#">GW170818</a>	$35.5^{+7.5}_{-4.7}$	$26.8^{+4.3}_{-5.2}$	$26.7^{+2.1}_{-1.7}$	$-0.09^{+0.18}_{-0.21}$	$59.8^{+4.8}_{-3.8}$	$0.67^{+0.07}_{-0.08}$	$2.7^{+0.5}_{-0.5}$	$3.4^{+0.5}_{-0.7} \times 10^{56}$	$1020^{+430}_{-360}$	$0.20^{+0.07}_{-0.07}$	39
<a href="#">GW170823</a>	$39.6^{+10.0}_{-6.6}$	$29.4^{+6.3}_{-7.1}$	$29.3^{+4.2}_{-3.2}$	$0.08^{+0.20}_{-0.22}$	$65.6^{+9.4}_{-6.6}$	$0.71^{+0.08}_{-0.10}$	$3.3^{+0.9}_{-0.8}$	$3.6^{+0.6}_{-0.9} \times 10^{56}$	$1850^{+840}_{-840}$	$0.34^{+0.13}_{-0.14}$	1651



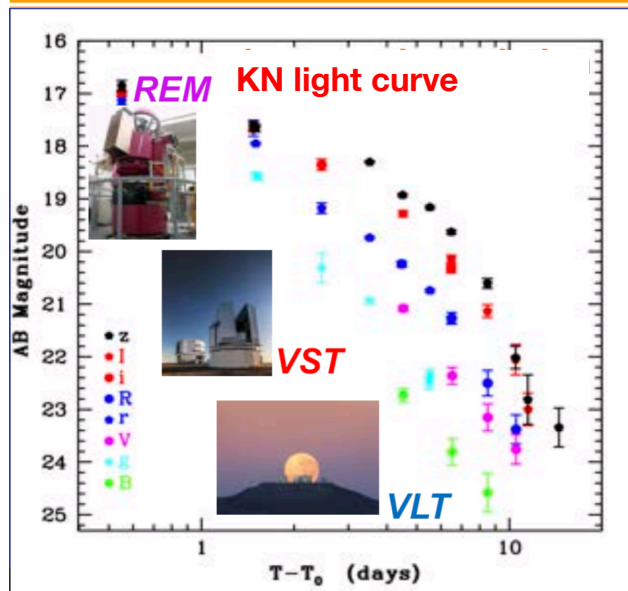


**Sept 2015 – Jan 2016: LVC O1 science run**  
**Nov 2016 – Aug 2017: LVC O2 science run**

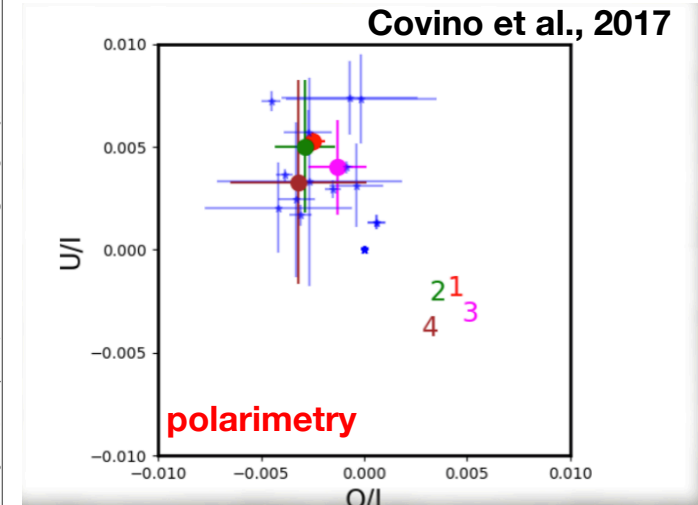
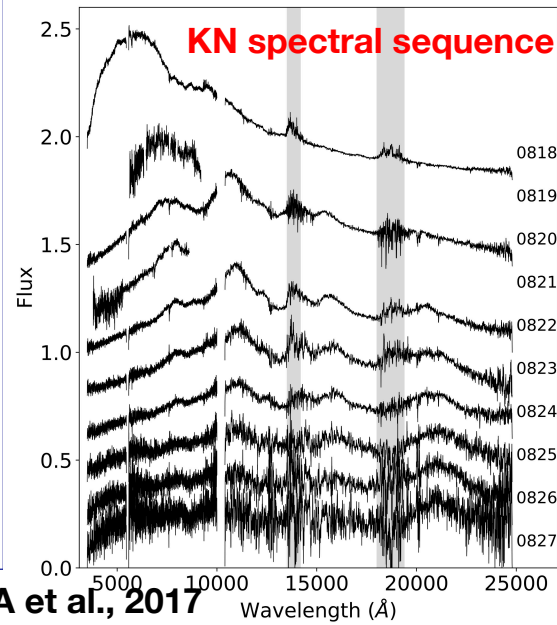


Event	$m_1/M_\odot$	$m_2/M_\odot$	$M/M_\odot$	$\chi_{\text{eff}}$	$M_f/M_\odot$	$a_f$	$E_{\text{rad}}/(M_\odot c^2)$	$\ell_{\text{peak}}/(\text{erg s}^{-1})$	$d_L/\text{Mpc}$	$z$	$\Delta\Omega/\text{deg}^2$
<a href="#">GW150914</a>	$35.6^{+4.8}_{-3.0}$	$30.6^{+3.0}_{-4.4}$	$28.6^{+1.6}_{-1.5}$	$-0.01^{+0.12}_{-0.13}$	$63.1^{+3.3}_{-3.0}$	$0.69^{+0.05}_{-0.04}$	$3.1^{+0.4}_{-0.4}$	$3.6^{+0.4}_{-0.4} \times 10^{56}$	$430^{+150}_{-170}$	$0.09^{+0.03}_{-0.03}$	180
<a href="#">GW151012</a>	$23.3^{+14.0}_{-5.5}$	$13.6^{+4.1}_{-4.8}$	$15.2^{+2.0}_{-1.1}$	$0.04^{+0.28}_{-0.19}$	$35.7^{+9.9}_{-3.8}$	$0.67^{+0.13}_{-0.11}$	$1.5^{+0.5}_{-0.5}$	$3.2^{+0.8}_{-1.7} \times 10^{56}$	$1060^{+540}_{-480}$	$0.21^{+0.09}_{-0.09}$	1555
<a href="#">GW151226</a>	$13.7^{+8.8}_{-3.2}$	$7.7^{+2.2}_{-2.6}$	$8.9^{+0.3}_{-0.3}$	$0.18^{+0.20}_{-0.12}$	$20.5^{+6.4}_{-1.5}$	$0.74^{+0.07}_{-0.05}$	$1.0^{+0.1}_{-0.2}$	$3.4^{+0.7}_{-1.7} \times 10^{56}$	$440^{+180}_{-190}$	$0.09^{+0.04}_{-0.04}$	1033
<a href="#">GW170104</a>	$31.0^{+7.2}_{-5.6}$	$20.1^{+4.9}_{-4.5}$	$21.5^{+2.1}_{-1.7}$	$-0.04^{+0.17}_{-0.20}$	$49.1^{+5.2}_{-3.9}$	$0.66^{+0.08}_{-0.10}$	$2.2^{+0.5}_{-0.5}$	$3.3^{+0.6}_{-0.9} \times 10^{56}$	$960^{+430}_{-410}$	$0.19^{+0.07}_{-0.08}$	924
<a href="#">GW170608</a>	$10.9^{+5.3}_{-1.7}$	$7.6^{+1.3}_{-2.1}$	$7.9^{+0.2}_{-0.2}$	$0.03^{+0.19}_{-0.07}$	$17.8^{+3.2}_{-0.7}$	$0.69^{+0.04}_{-0.04}$	$0.9^{+0.05}_{-0.1}$	$3.5^{+0.4}_{-1.3} \times 10^{56}$	$320^{+120}_{-110}$	$0.07^{+0.02}_{-0.02}$	396
<a href="#">GW170729</a>	$50.6^{+16.6}_{-10.2}$	$34.3^{+9.1}_{-10.1}$	$35.7^{+6.5}_{-4.7}$	$0.36^{+0.21}_{-0.25}$	$80.3^{+14.6}_{-10.2}$	$0.81^{+0.07}_{-0.13}$	$4.8^{+1.7}_{-1.7}$	$4.2^{+0.9}_{-1.5} \times 10^{56}$	$2750^{+1350}_{-1320}$	$0.48^{+0.19}_{-0.20}$	1033
<a href="#">GW170809</a>	$35.2^{+8.3}_{-6.0}$	$23.8^{+5.2}_{-5.1}$	$25.0^{+2.1}_{-1.6}$	$0.07^{+0.16}_{-0.16}$	$56.4^{+5.2}_{-3.7}$	$0.70^{+0.08}_{-0.09}$	$2.7^{+0.6}_{-0.6}$	$3.5^{+0.6}_{-0.9} \times 10^{56}$	$990^{+320}_{-380}$	$0.20^{+0.05}_{-0.07}$	340
<a href="#">GW170814</a>	$30.7^{+5.7}_{-3.0}$	$25.3^{+2.9}_{-4.1}$	$24.2^{+1.4}_{-1.1}$	$0.07^{+0.12}_{-0.11}$	$53.4^{+3.2}_{-2.4}$	$0.72^{+0.07}_{-0.05}$	$2.7^{+0.4}_{-0.3}$	$3.7^{+0.4}_{-0.5} \times 10^{56}$	$580^{+160}_{-210}$	$0.12^{+0.03}_{-0.04}$	87
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<a href="#">GW170823</a>	$39.6^{+10.0}_{-6.6}$	$29.4^{+6.3}_{-7.1}$	$29.3^{+4.2}_{-3.2}$	$0.08^{+0.20}_{-0.22}$	$65.6^{+9.4}_{-6.6}$	$0.71^{+0.08}_{-0.10}$	$3.3^{+0.9}_{-0.8}$	$3.6^{+0.6}_{-0.9} \times 10^{56}$	$1850^{+840}_{-840}$	$0.34^{+0.13}_{-0.14}$	1651

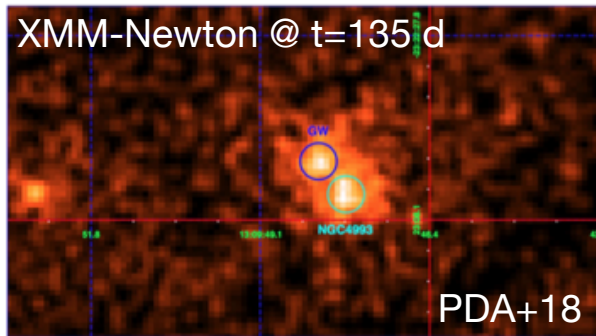




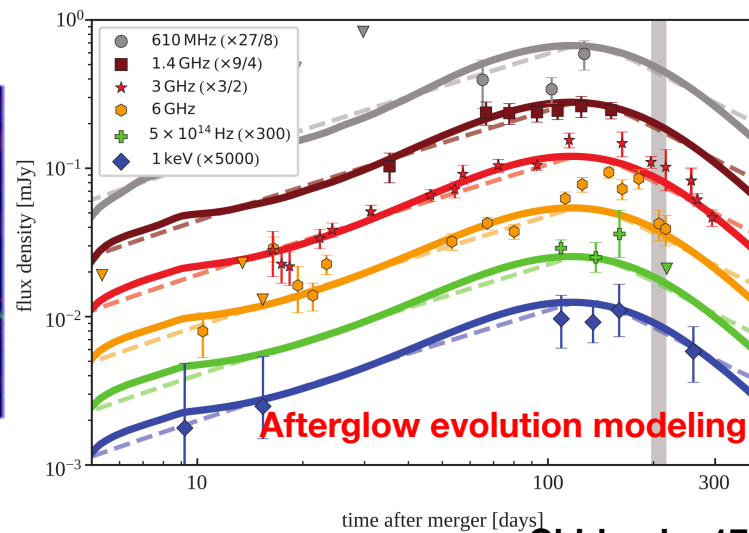
Pian, PDA et al., 2017



Full characterization of the KN and GRB properties

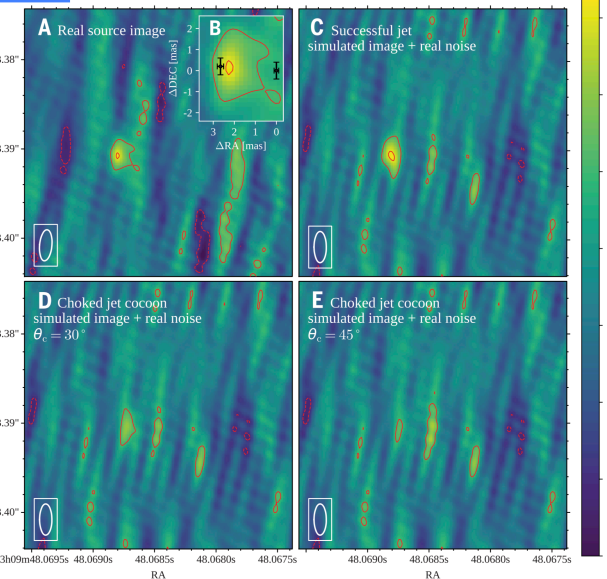


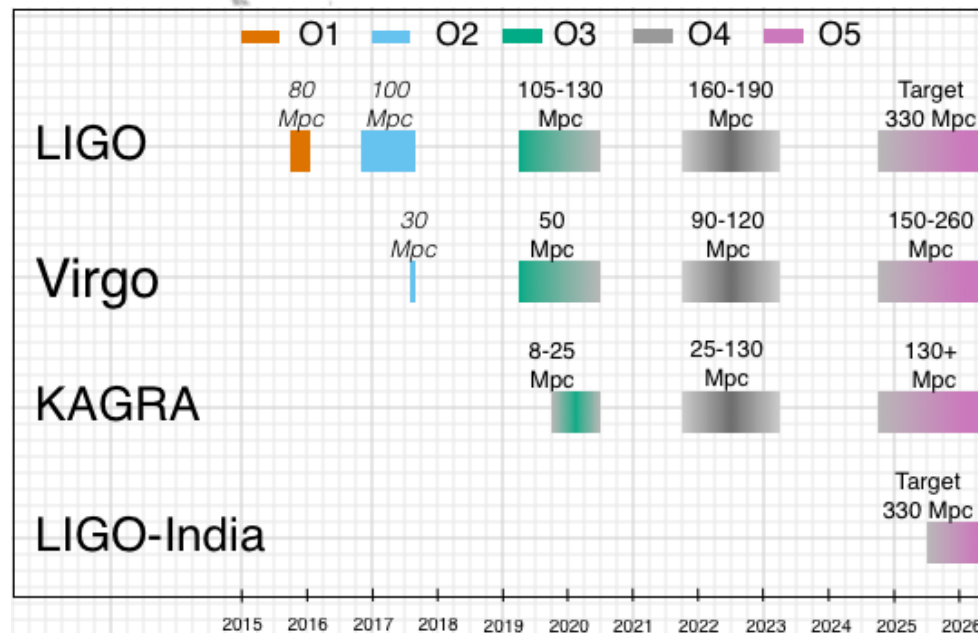
detection of the afterglow at the peak



Ghirlanda+17

Evidence for a structured GRB jet





## Search (& follow-up) European teams



**Governing Council:** M. Branchesi, E. Brocato, P. D'Avanzo, J. Hjorth, P. Jonker, E. Pian, S. Smartt (Chair), J. Sollerman, D. Steeghs, N. Tanvir.  
**Executive Committee:** A. Levan (Chair), M. Fraser, K. Maguire, D. Malesani, O. S. Salafia, S. Vergani.



### A collaboration of ~ 200 ESO scientists

Approved program during Oct 2018 – Mar 2020 fully covering O3. Time for EM counterparts **follow-up** on every useful **VLT** instrument + **ALMA** **HST** and **JWST**.



**VINROUGE**

[www.star.le.ac.uk/nrt3/VINROUGE/](http://www.star.le.ac.uk/nrt3/VINROUGE/)



[www.grawita.inaf.it](http://www.grawita.inaf.it)

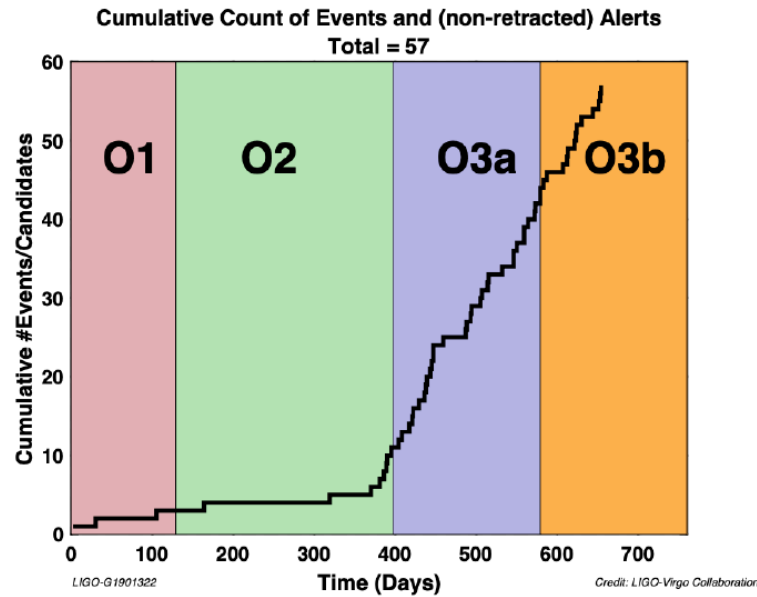
The **GW** Optical  
Transient Observer

**GOTO**

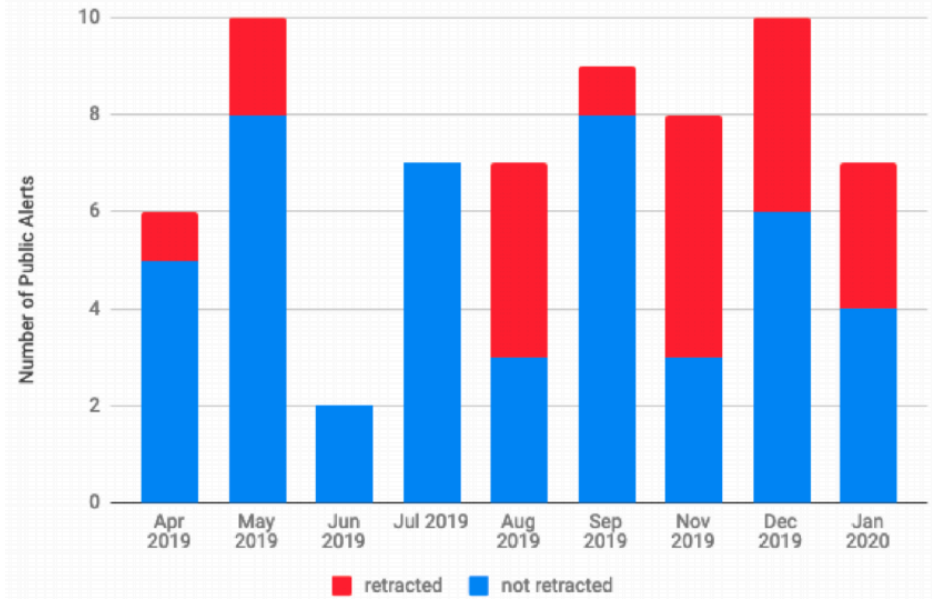
[www.goto-observatory.org](http://www.goto-observatory.org)



[www.pessto.org](http://www.pessto.org)



**O3 Public Alerts by Month**







## GW190814: Gravitational Waves from the Coalescence of a 23 Solar Mass Black Hole with a 2.6 Solar Mass Compact Object

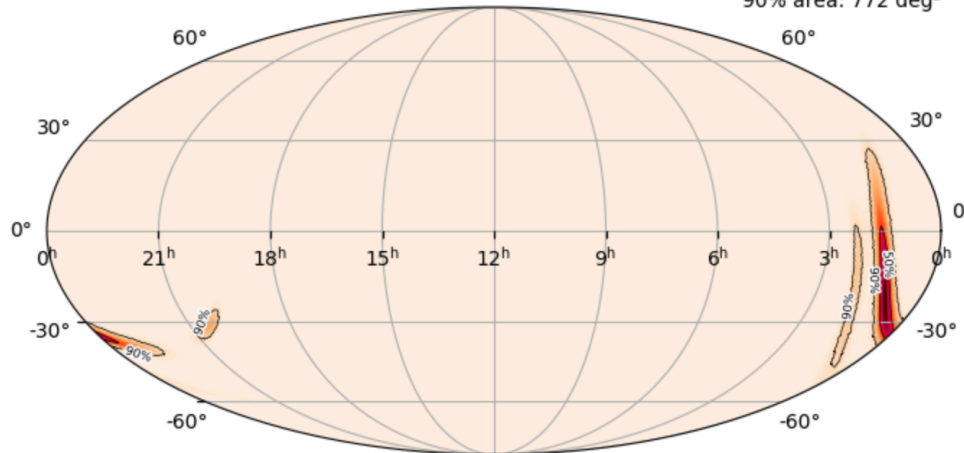
R. Abbott<sup>1</sup>, T. D. Abbott<sup>2</sup>, S. Abraham<sup>3</sup>, F. Acernese<sup>4,5</sup>, K. Ackley<sup>6</sup>, C. Adams<sup>7</sup>, R. X. Adhikari<sup>1</sup>, V. B. Adya<sup>8</sup>, C. Affeldt<sup>9,10</sup>, M. Agathos<sup>11,12</sup>, K. Agatsuma<sup>13</sup>, N. Aggarwal<sup>14</sup>, O. D. Aguiar<sup>15</sup>, A. Aich<sup>16</sup>, L. Aiello<sup>17,18</sup>, A. Ain<sup>3</sup>, P. Ajith<sup>19</sup>, S. Akcay<sup>11,20</sup>,

### Abstract

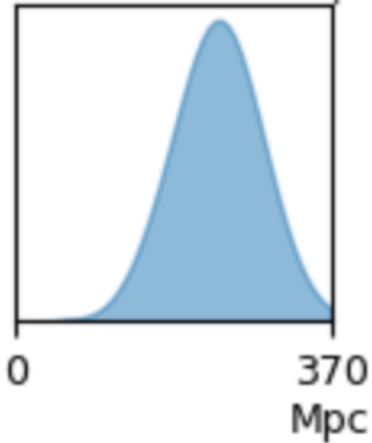
We report the observation of a compact binary coalescence involving a  $22.2\text{--}24.3 M_{\odot}$  black hole and a compact object with a mass of  $2.50\text{--}2.67 M_{\odot}$  (all measurements quoted at the 90% credible level). The gravitational-wave signal, GW190814, was observed during LIGO's and Virgo's third observing run on 2019 August 14 at 21:10:39 UTC and has a signal-to-noise ratio of 25 in the three-detector network. The source was localized to  $18.5 \text{ deg}^2$  at a distance of  $241_{-45}^{+41}$  Mpc; no electromagnetic counterpart has been confirmed to date. The source has the most unequal mass ratio yet measured with gravitational waves,  $0.112_{-0.009}^{+0.008}$ , and its secondary component is either the lightest black hole or the heaviest neutron star ever discovered in a double compact-object system. The dimensionless spin of the primary black hole is tightly constrained to  $\leq 0.07$ . Tests of general relativity reveal no measurable deviations from the theory, and its prediction of higher-multipole emission is confirmed at high confidence. We estimate a merger rate density of  $1\text{--}23 \text{ Gpc}^{-3} \text{ yr}^{-1}$  for the new class of binary coalescence sources that GW190814 represents. Astrophysical models predict that binaries with mass ratios similar to this event can form through several channels, but are unlikely to have formed in globular clusters. However, the combination of mass ratio, component masses, and the inferred merger rate for this event challenges all current models of the formation and mass distribution of compact-object binaries.

21:10:39 UT

event ID: G347292  
50% area: 133 deg<sup>2</sup>  
90% area: 772 deg<sup>2</sup>

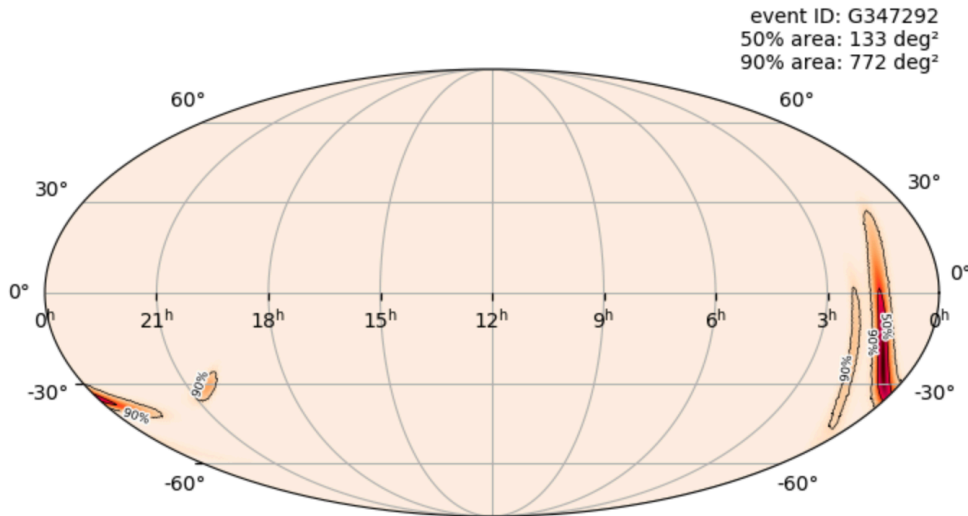


event ID: G347292  
distance:  $236 \pm 53$  Mpc

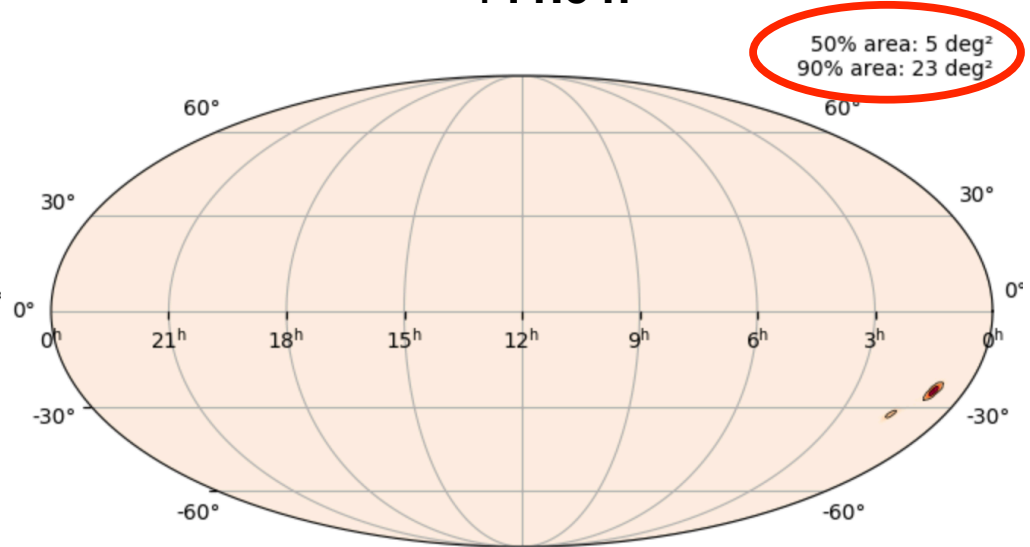


MassGap	100%
Terrestrial	<1%
NSBH	0%
BNS	0%
BBH	0%

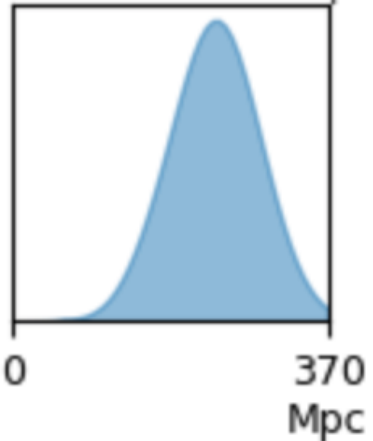
21:10:39 UT



+11.6 h

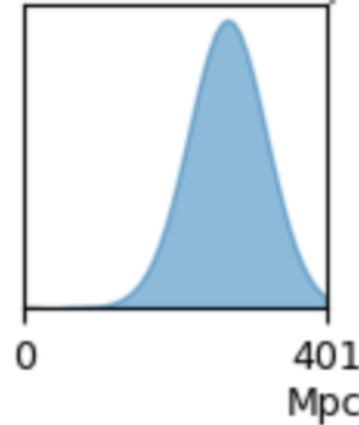


event ID: G347292  
distance: 236±53 Mpc

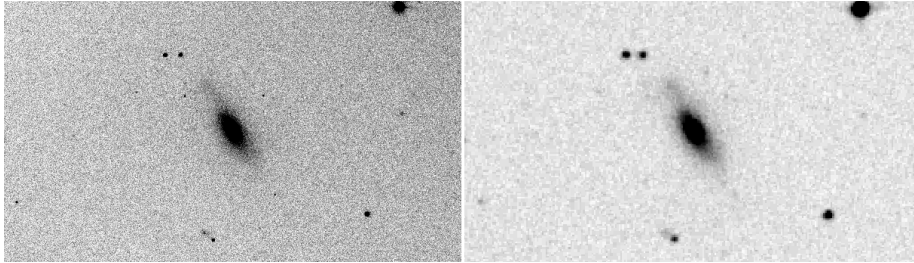


MassGap	100%
Terrestrial	<1%
NSBH	0%
BNS	0%
BBH	0%

distance: 267±52 Mpc



NSBH	>99%
MassGap	<1%
Terrestrial	0%
BNS	0%
BBH	0%



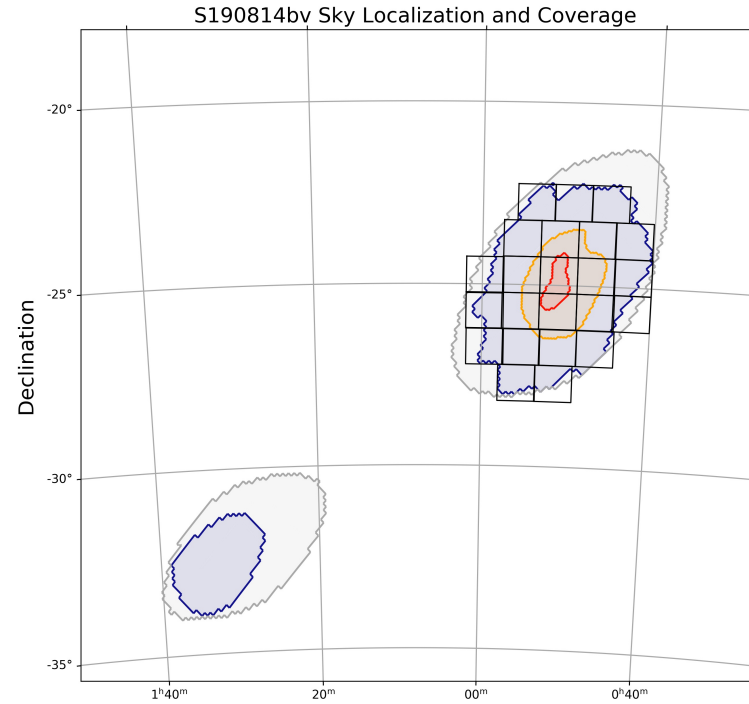
GTC, VLT, LBT, WHT, TNG, NOT, LT,  
GROND coordinated  
observations of **more than 70 galaxies**,  
over multiple epochs within the 90%  
localization of the GW

+

Wide-field observations  
(VST, PS, ATLAS, GOTO)

**No e.m. counterpart found**

**Ackley+20**  
**(ENGRAVE collaboration)**



courtesy: G. Greco

### 5 VST epochs

$\Delta t$	r AB mag	Probability
+11.5 h	20.9	60.7%
+1.5 d	21.9	71.5%
+4.3 d	21.7	87.7%
+7.2 d	21.8	87.7%
+14.5 d	22.0	87.7%

- During **O1** (Sep 2015 – Jan 2016) and **O2** (Nov 2016 - Aug 2017) we carried out observational campaigns aimed at the search of e.m. counterparts and follow-up of promising e.m. candidate counterparts of **7 GW triggers (including the successful GW 170817 campaign)**.



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- During **O3** (Apr 2019 – Mar 2020) we carried out observational campaigns aimed at the search of e.m. counterparts and follow-up of promising e.m. candidate counterparts of **13 GW triggers**.

- **After the end of O3** (due to the COVID pandemic) we focused our observational activities on the follow-up of KNe candidates found in surveys and of well-localised ( $< 50 \text{ deg}^2$ ) short GRBs found by Fermi/GBM. On this topic, we observed **2 KN candidates, 1 short GRB, 1 candidate off-axis afterglow of a short GRB**.



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Since 2015 we published:

- **64** GCN Circulars, ATel, Astronote
- **8** refereed papers (546 cit., 68 cit./paper) -> GRAWITA-led
- **3** refereed paper (2132 cit., 711 cit./paper) -> GRAWITA participation
- **32** refereed paper (2258 cit, 71 cit./paper) -> GW/KN related papers with participation of (or led by) GRAWITA scientists
- of the above, **9** circulars and **6** papers are based on **TNG** data (274 cit., 46 cit./paper)



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- We have been awarded of **435** hours of observing time at TNG over **15** semesters (**29 h/semester** on average). Current LTP will run up to Feb 2023.

		Area (deg <sup>2</sup> ) 90% c.r.	Area (deg <sup>2</sup> ) 90% c.r.	Area (deg <sup>2</sup> ) 90% c.r.
O3	HLV	270 <sup>+34</sup> <sub>-20</sub>	330 <sup>+24</sup> <sub>-31</sub>	280 <sup>+30</sup> <sub>-23</sub>
→ O4	HLVK	33 <sup>+5</sup> <sub>-5</sub>	50 <sup>+8</sup> <sub>-8</sub>	41 <sup>+7</sup> <sub>-6</sub>
		Comoving Volume (10 <sup>3</sup> Mpc <sup>3</sup> ) 90% c.r.	Comoving Volume (10 <sup>3</sup> Mpc <sup>3</sup> ) 90% c.r.	Comoving Volume (10 <sup>3</sup> Mpc <sup>3</sup> ) 90% c.r.
O3	HLV	120 <sup>+19</sup> <sub>-24</sub>	860 <sup>+150</sup> <sub>-150</sub>	16000 <sup>+2200</sup> <sub>-2500</sub>
→ O4	HLVK	52 <sup>+10</sup> <sub>-9</sub>	430 <sup>+100</sup> <sub>-78</sub>	7700 <sup>+1500</sup> <sub>-920</sub>
Observation Run	Network	Expected BNS Detections	Expected NSBH Detections	Expected BBH Detections
O3	HLV	1 <sup>+12</sup> <sub>-1</sub>	0 <sup>+19</sup> <sub>-0</sub>	17 <sup>+22</sup> <sub>-11</sub>
→ O4	HLVK	10 <sup>+52</sup> <sub>-10</sub>	1 <sup>+91</sup> <sub>-1</sub>	79 <sup>+89</sup> <sub>-44</sub>

**We are at the dawn of a new, exciting, promising, era for astrophysics. No doubt that there is a lot of attention, efforts, planning, expectations from the international community.**

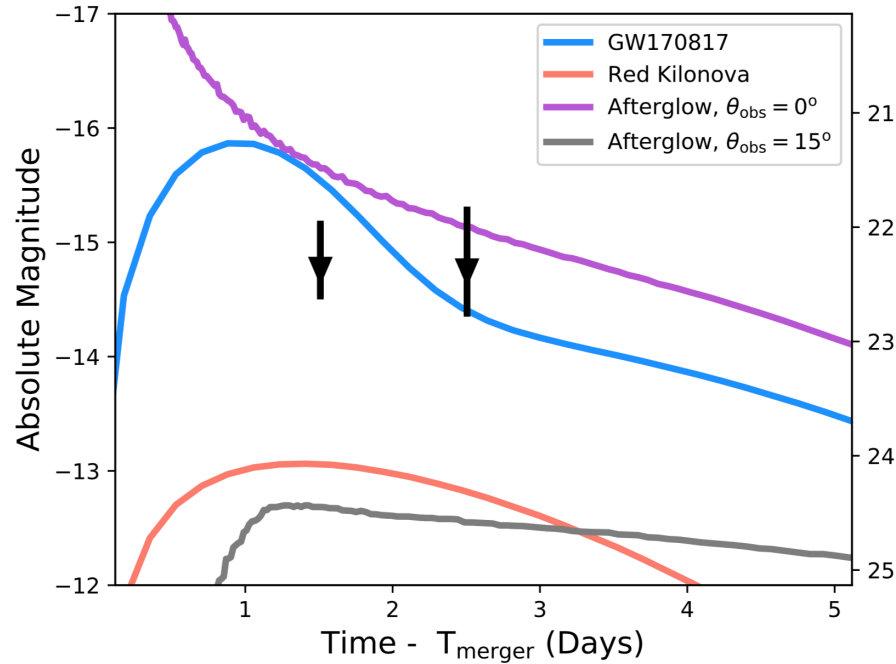
**Basically all the main forthcoming facilities have multi-messenger astrophysics in their science case (or will play anyway a role): VRO (LSST), CTA, SKA, JWST, e-ELT. We (GRAWITA and INAF more in general) are involved in all of them.**

**The TNG characteristics and the diversity of its instrumentation represent a key asset in this context. Working in collaborations and consortia having access to different facilities (all sizes, all wavelengths) ensure an optimal use of the telescope.**

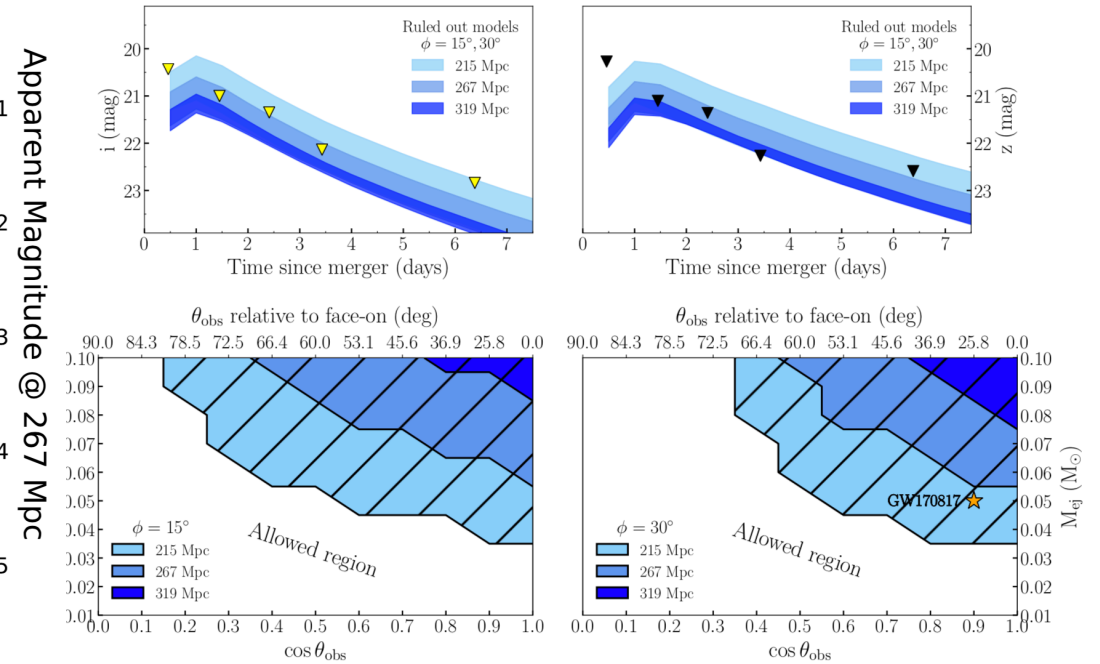
**Great collaboration and interaction with TNG Director and Staff  
(davvero, grazie ragazzi!)**



### Gomez+19



### Andreoni+19



**Strong constraints if the lower distance of S190814bv is assumed**