

Exploring the Universe with Gamma-Ray Burst Afterglows

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on behalf of the CIBO collaboration

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Hotel Taburiente a Los Cancajos, 19-21 October 2021



Coordinamento Italiano Burst Ottici

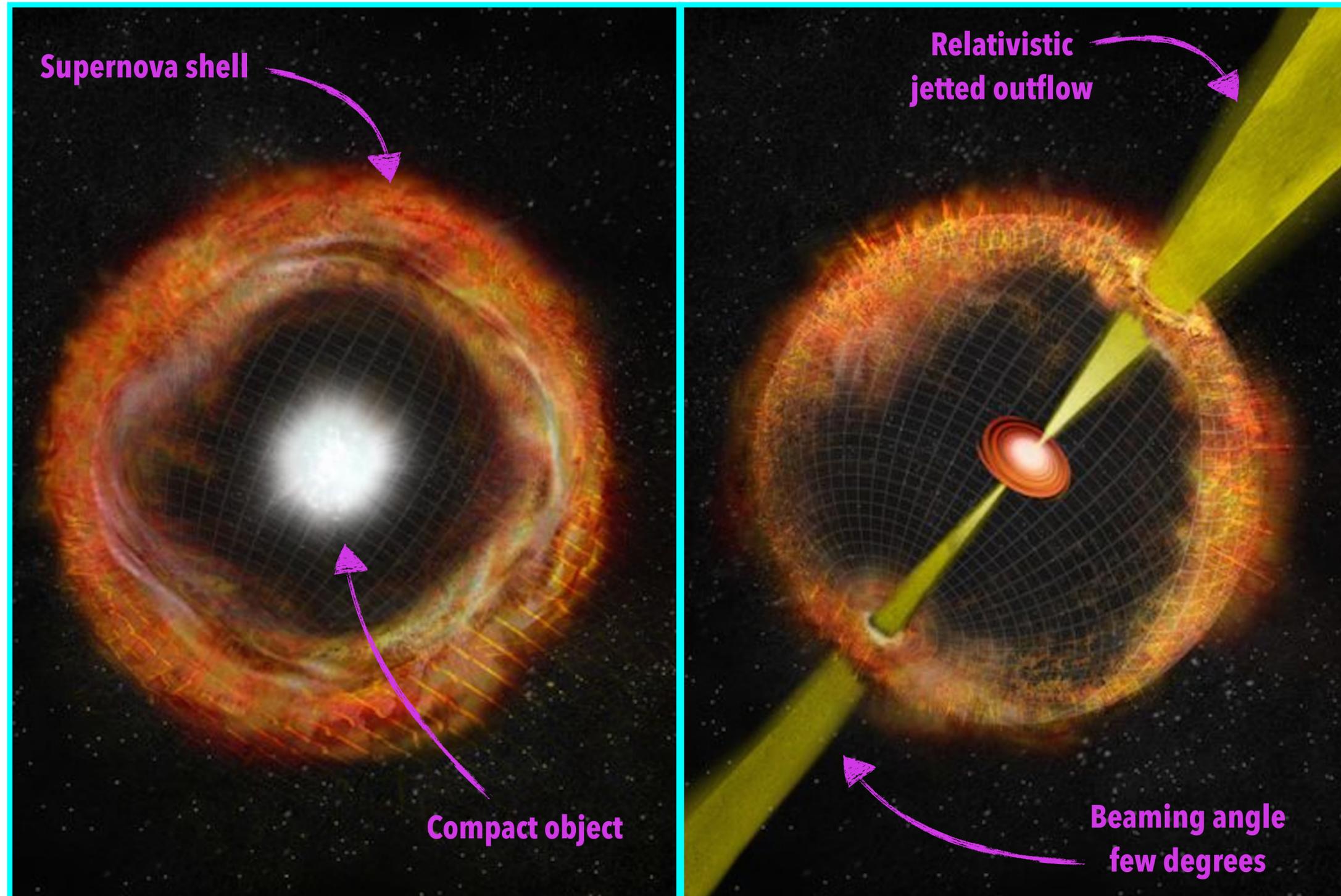
The **C.I.B.O.** collaboration, formed in 2000, involves most of the Italian astronomers interested in optical and infrared observations of the GRB afterglows and their host galaxies (HGs). Many members of the collaboration are also part of the Italian *Swift* team.

C.I.B.O. activities:

- Follow-up observations of GRBs & HGs with **TNG** since 2000
- Follow-up observations of GRBs with **small Italian Telescopes** since 2000
- Follow-up observations of GRBs with **REM** since 2004
- Follow up observations of GRBs & HGs with **LBT** since 2008

Gamma-Ray Bursts

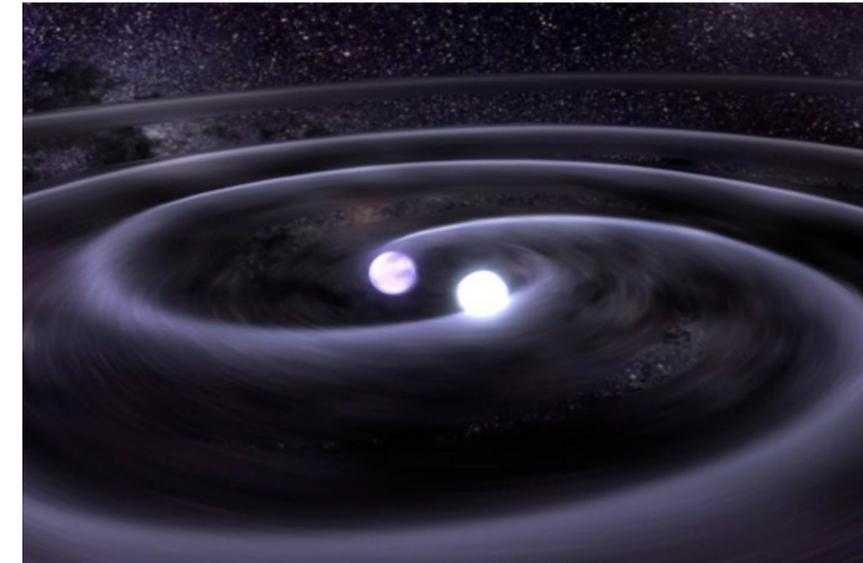
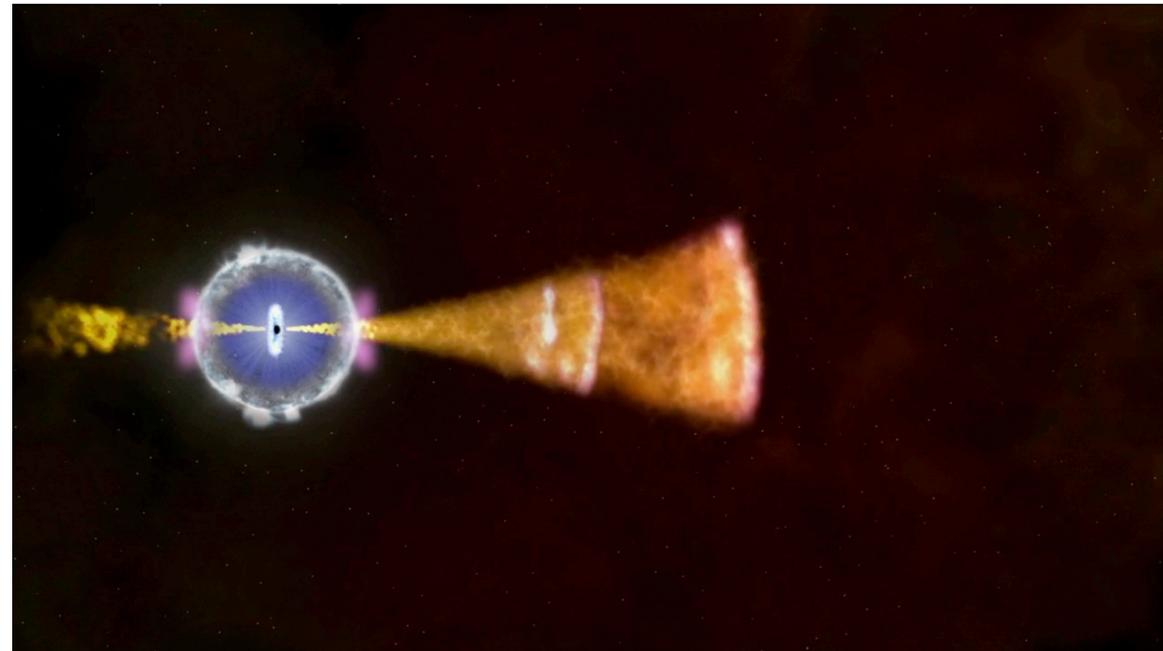
Brief intense (short or long) pulses of gamma-rays - Most energetic events (in the EM regime) in the Universe



Science with Gamma-Ray Bursts

- GRBs Physics

1. Shocks
2. The role of magnetic field
3. Jets
4. Accretion/ejection: extreme regimes

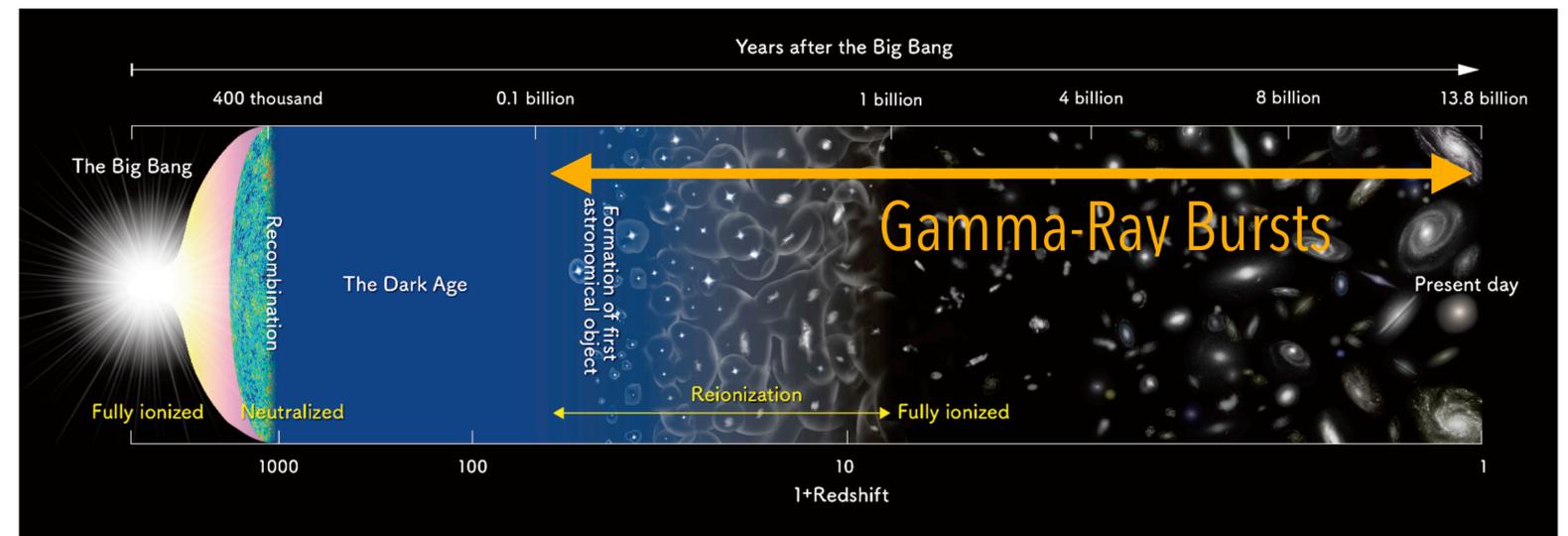


- Progenitors

1. Long GRBs: GRB-SN connection
2. Short GRBs: compact objects merging (GW)

- Long GRBs as cosmological probes

1. Connection with massive star formation
2. From the local Universe to the re-ionisation era
3. Circumburst environment / IGM
4. Chemical history of the Universe



Since 2004 *Swift* observed 1499 GRBs: legacy/statistical approach to tackle the above science cases

Gamma-Ray Bursts with TNG: a legacy approach

Since AOT22 (2010) we started a series of Long Term Programs focused on well defined scientific topics
(ensure a proper follow-up of rare but flagship events):

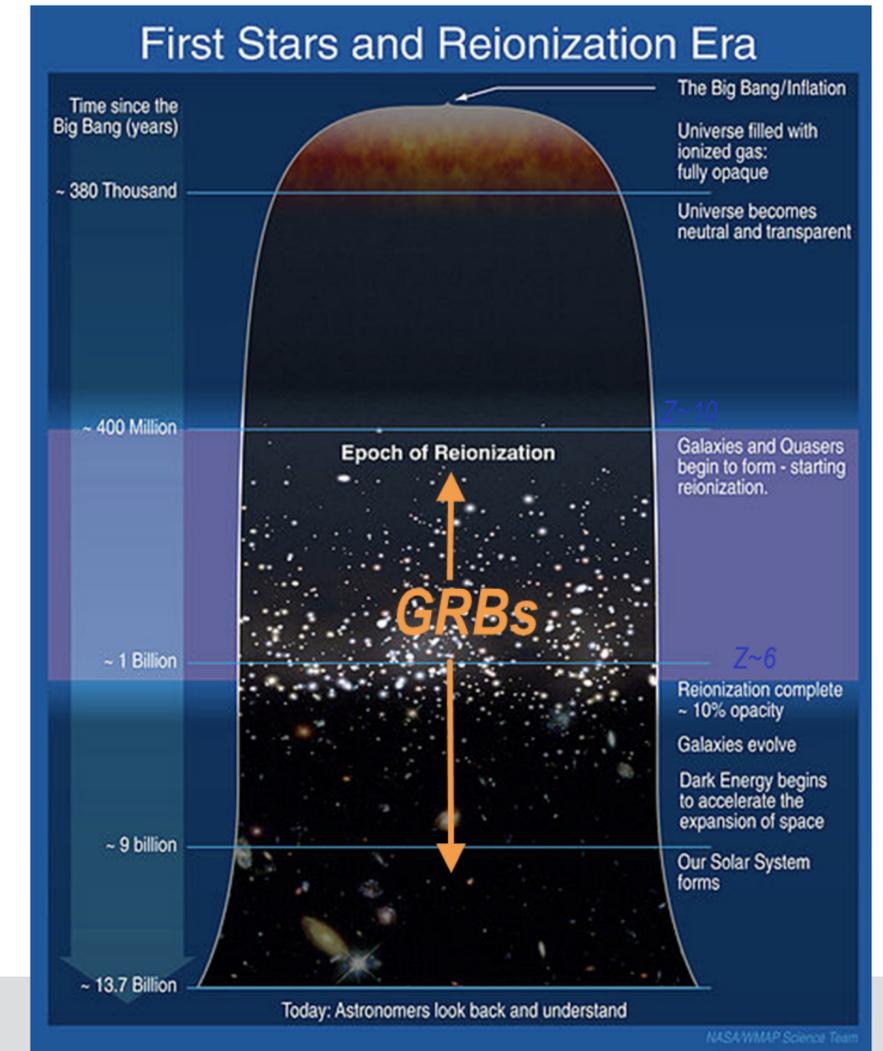
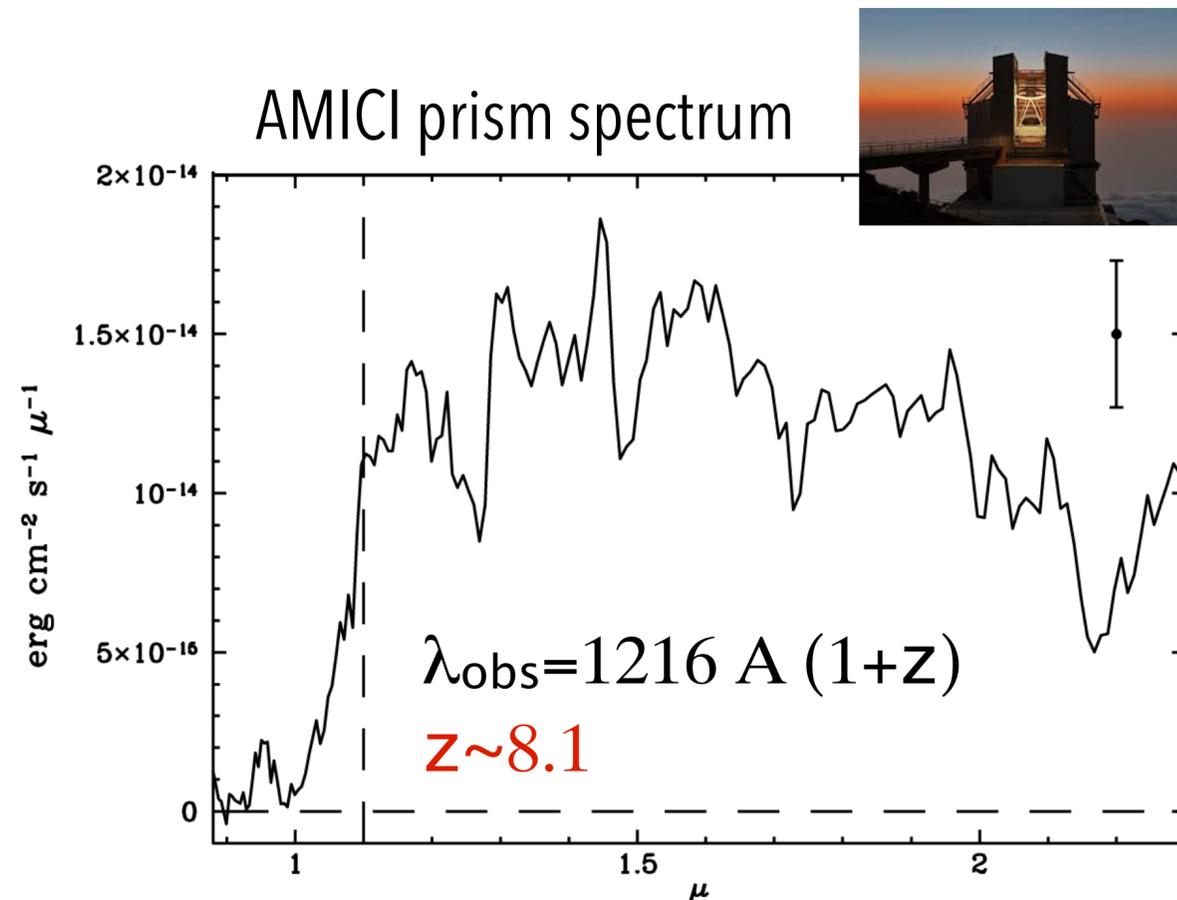
- candidate high redshift GRBs
- GRB–SN association
- short GRBs (strong connection with GW events!!)
- events belonging to complete (flux–limited) samples (bright GRBs)

MANDATORY NOT TO MISS

GRBs as cosmic probes: GRB090423

Long GRBs are detectable from the local Universe to very high redshifts:

- cosmic star formation history
- metallicity & dust evolution
- properties of faint galaxies (missed by 'traditional' surveys)



nature

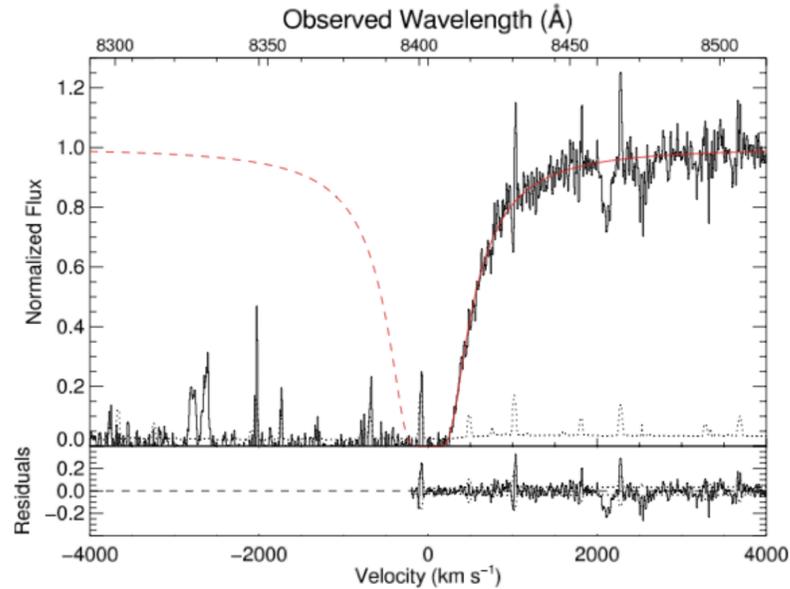
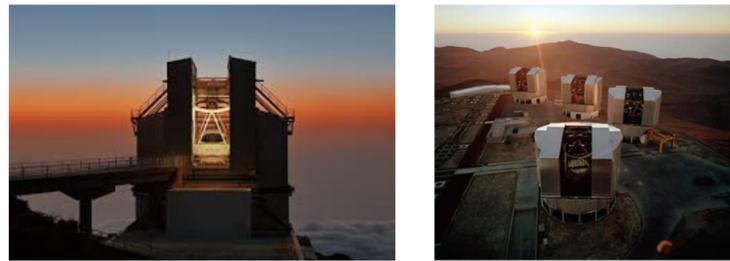
LETTERS

GRB 090423 at a redshift of $z \approx 8.1$

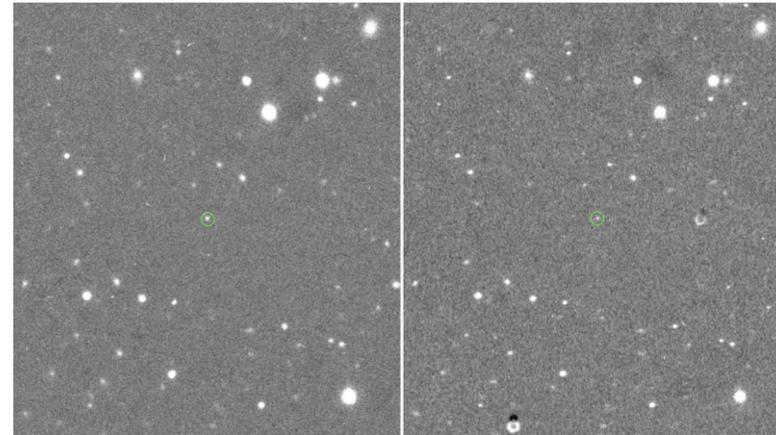
R. Salvaterra¹, M. Della Valle^{2,3,4}, S. Campana¹, G. Chincarini^{1,5}, S. Covino¹, P. D'Avanzo^{1,5}, A. Fernández-Soto⁶, C. Guidorzi⁷, F. Mannucci⁸, R. Margutti^{1,5}, C. C. Thöne¹, L. A. Antonelli⁹, S. D. Barthelmy¹⁰, M. De Pasquale¹¹, V. D'Elia⁹, F. Fiore⁹, D. Fugazza¹, L. K. Hunt⁸, E. Maiorano¹², S. Marinoni^{13,14}, F. E. Marshall¹⁰, E. Molinari^{1,13}, J. Nousek¹⁵, E. Pian^{16,17}, J. L. Racusin¹⁵, L. Stella⁹, L. Amati¹², G. Andreuzzi¹³, G. Cusumano¹⁸, E. E. Fenimore¹⁹, P. Ferrero²⁰, P. Giommi²¹, D. Guetta⁹, S. T. Holland^{10,22,23}, K. Hurley²⁴, G. L. Israel⁹, J. Mao¹, C. B. Markwardt^{10,23,25}, N. Masetti¹², C. Pagani¹⁵, E. Palazzi¹², D. M. Palmer¹⁸, S. Piranomonte⁹, G. Tagliaferri¹ & V. Testa⁹

GRBs as cosmic probes: rareness (17 events with $z > 5$)

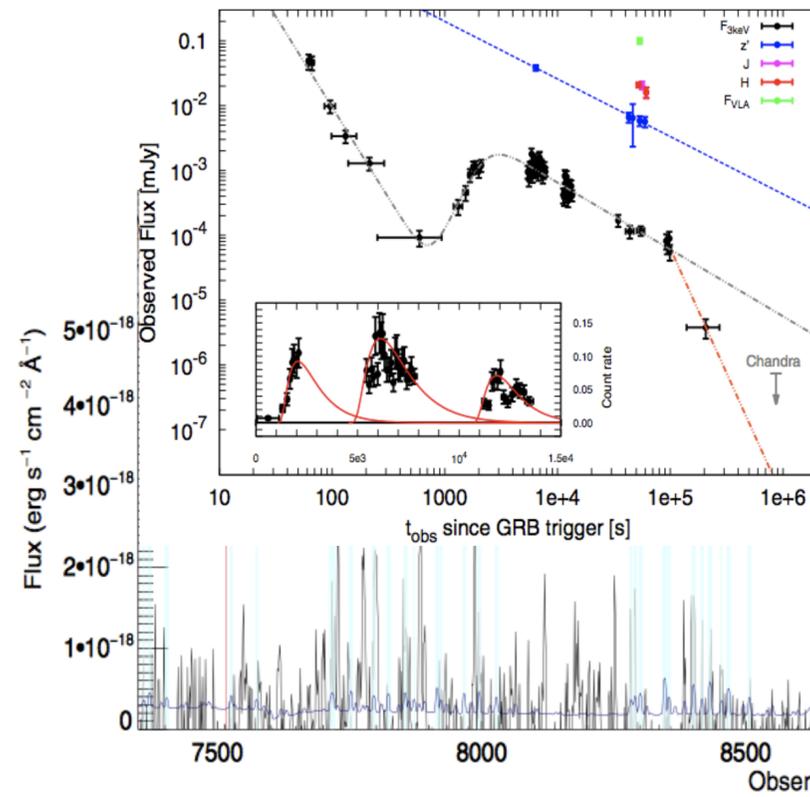
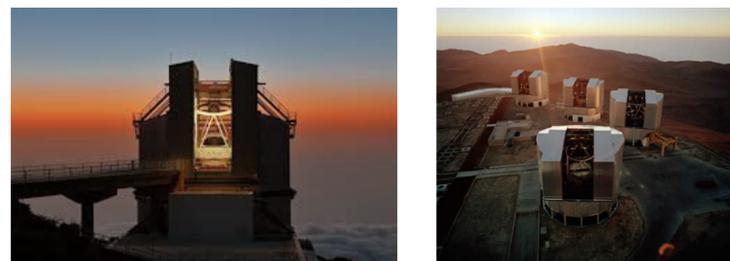
GRB 130606A ($z = 5.913$)
Hartoog et al. 2015



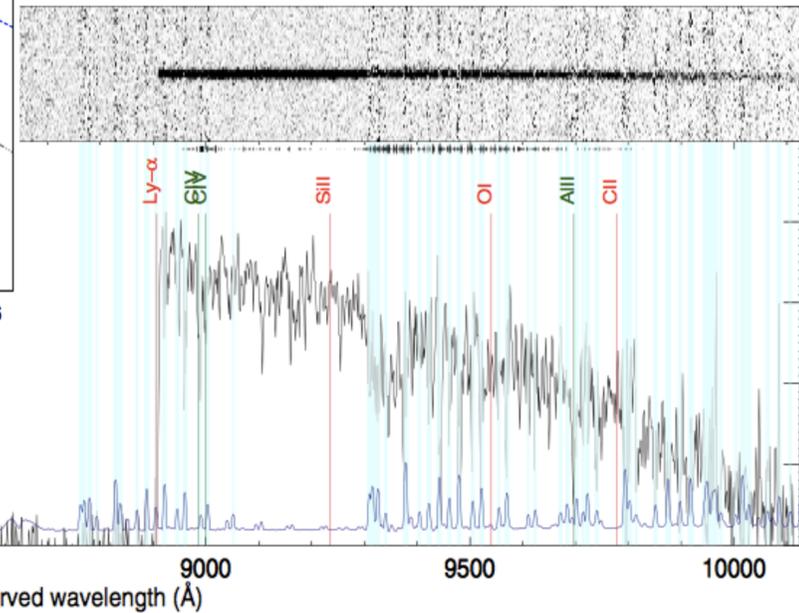
TNG z-band



GRB 140515A ($z = 6.32$)
Melandri et al. 2015



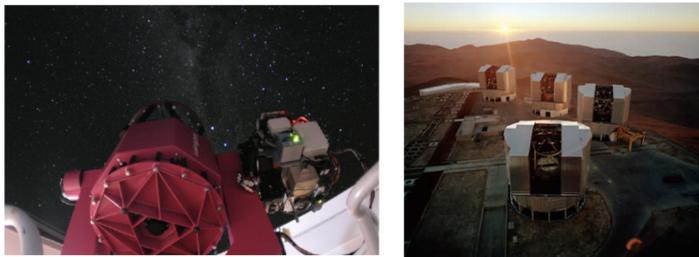
TNG JH-bands



GRB	Redshift
050904	6.29
080913	6.695
090423	8.2
090429B	9.4ph
100905A	7.9ph
120521C	6.0ph
120923A	7.8
140515A	6.32
210905A	6.318

GRBs as cosmic probes: rareness

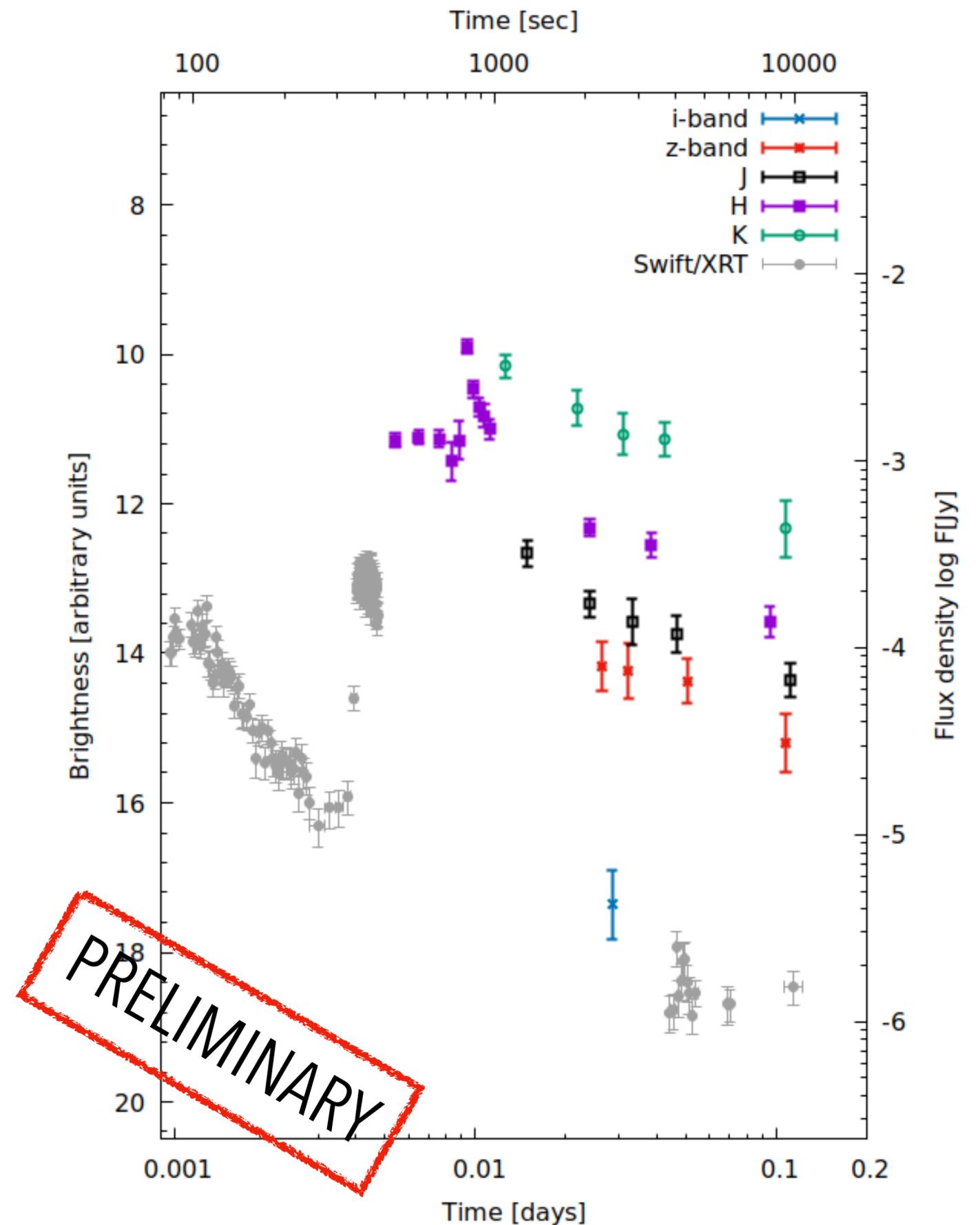
GRB 210905A ($z = 6.318$)
Rossi et al. 2021 in preparation



REM *izJHK* light curve
 continuous observations for the
 first 2 hr after the GRB event

GRB	Redshift
060522	5.11
050502B	5.2 ph
140304A	5.283
050814	5.3
131227A	5.3
060927	5.467
201221A	5.7
130606A	5.913
120521C	6.0 ph
050904	6.295
210905A	6.318
140515A	6.327
080913	6.695
100905A	7.9 ph
090423	8.26
120923A	7.8
090429B	9.4 ph

only 12 (5) events
with $z_{\text{spec}} > 5$ (6)

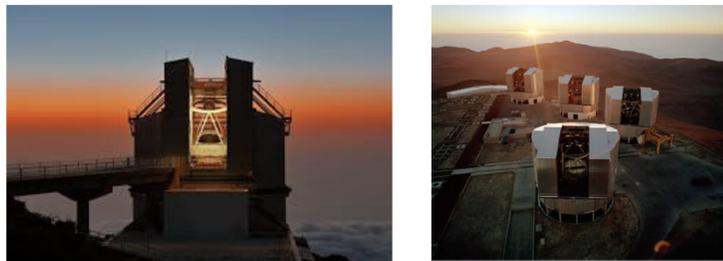


GRB-SN connection (observed for about 20 events)

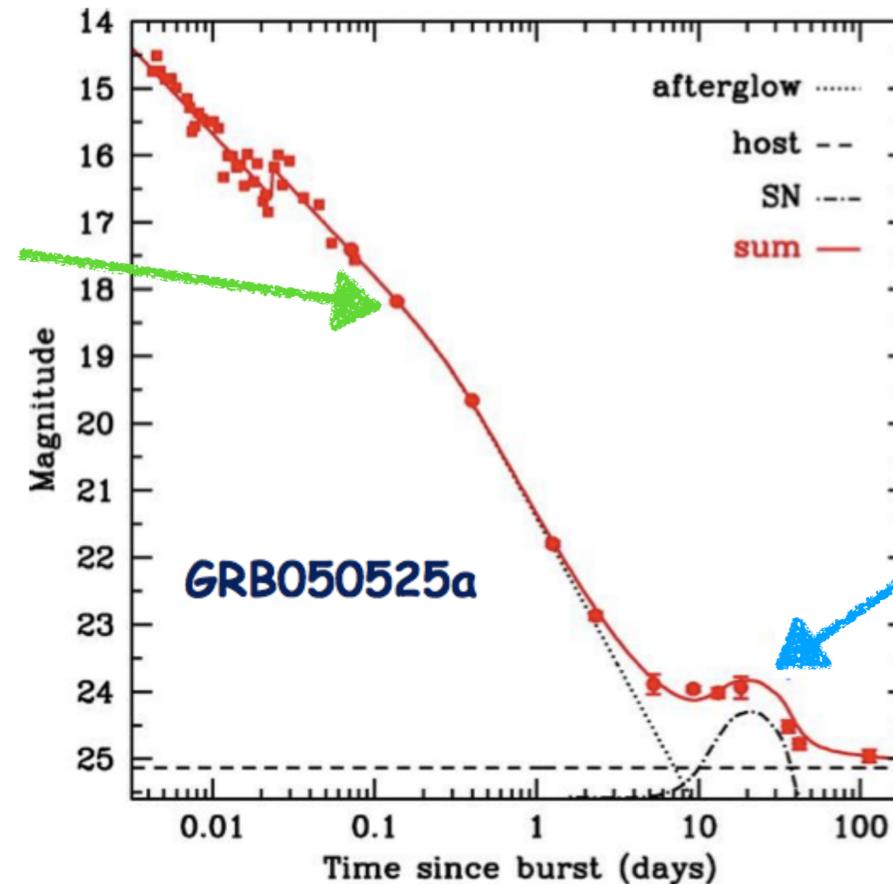
Nearby Long GRBs are usually associated with broad lined Type Ib-c SN:

- powering mechanism and progenitor features
- properties of the circumstellar medium and host-galaxies

GRB 050525A/SN2005nc ($z = 0.606$)
Della Valle et al. 2006



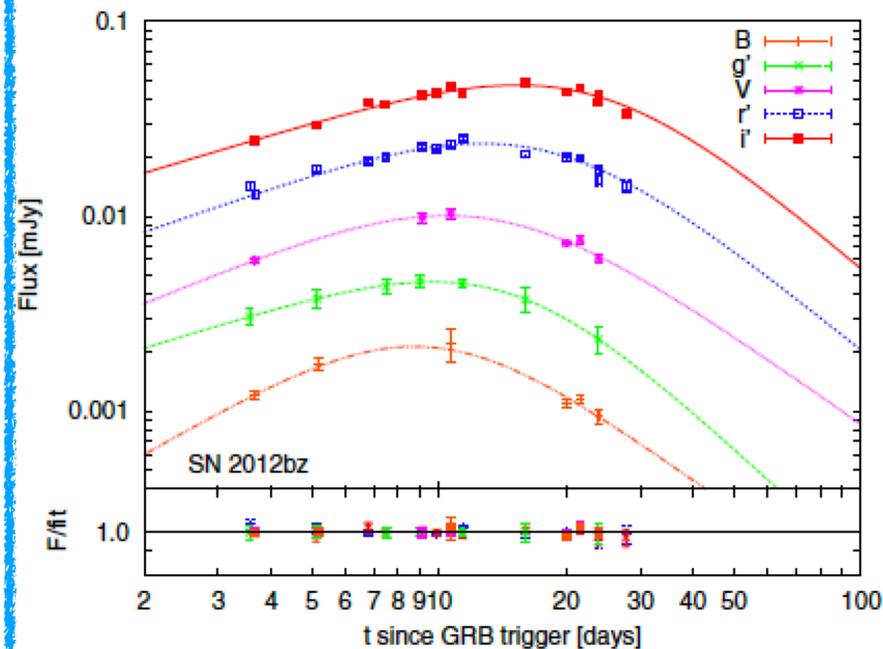
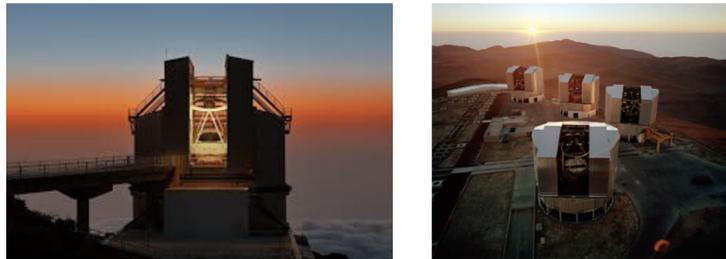
TNG optical light curve



VLT spectrum near maximum light

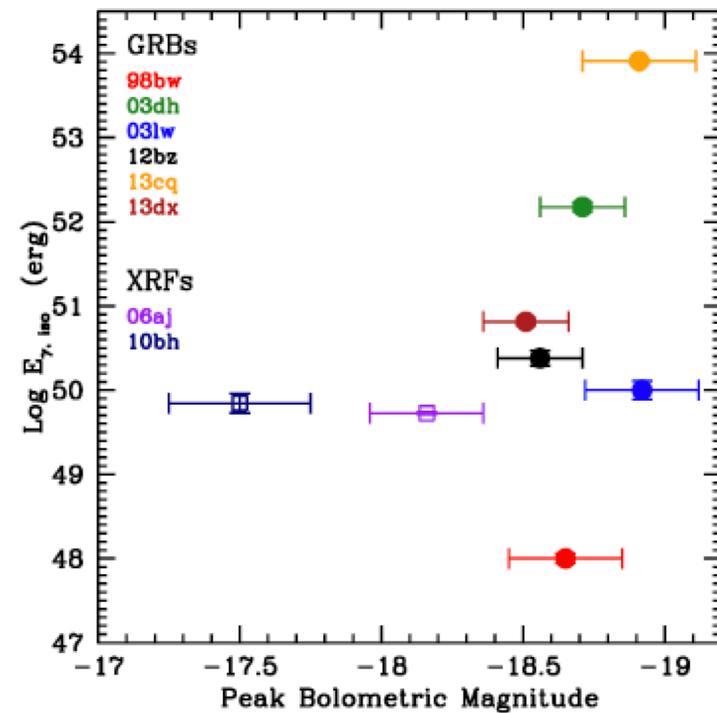
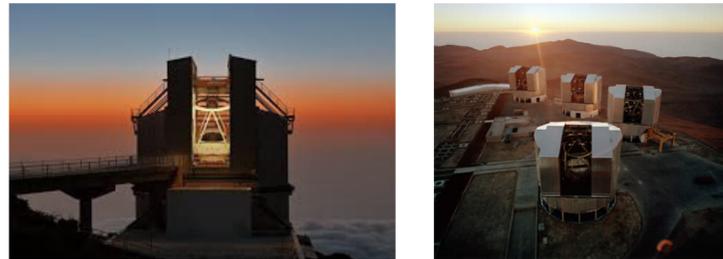
GRB-SN connection: ordinary events, ordinary monsters and real monsters

GRB 120422A/SN2012bz ($z = 0.283$)
Melandri et al. 2012



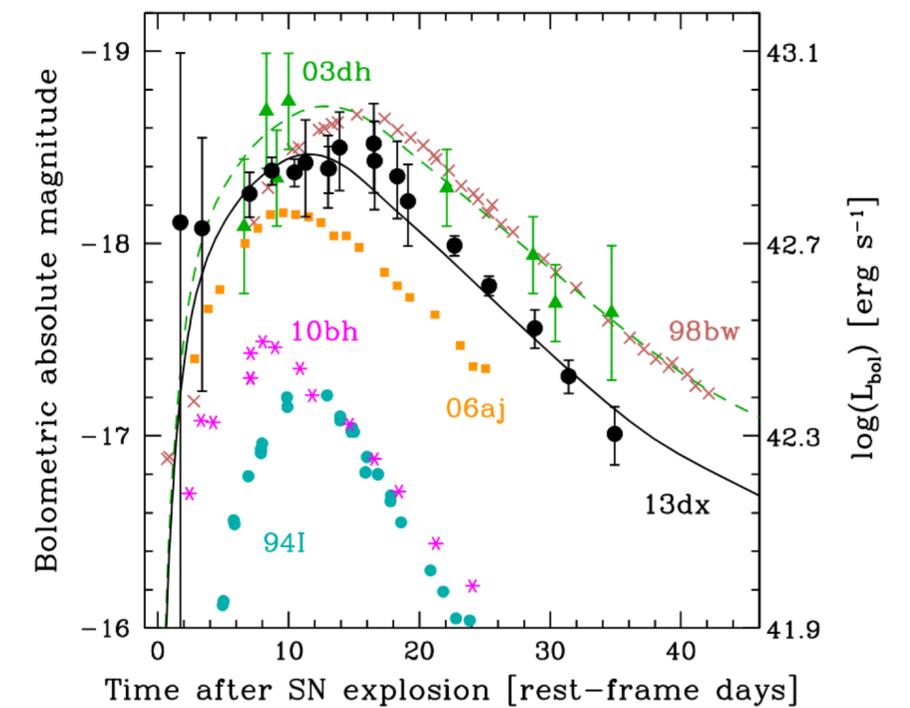
TNG *griz* obs.

GRB 130427A/SN2013cq ($z = 0.3399$)
Melandri et al. 2014



TNG *gri* obs.

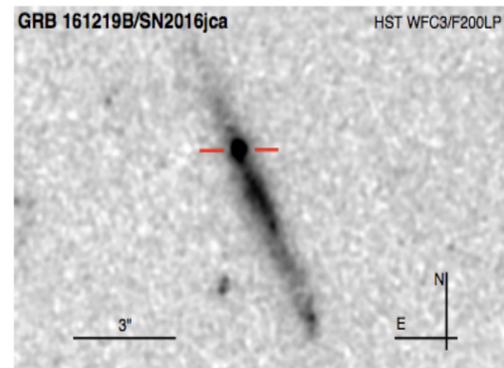
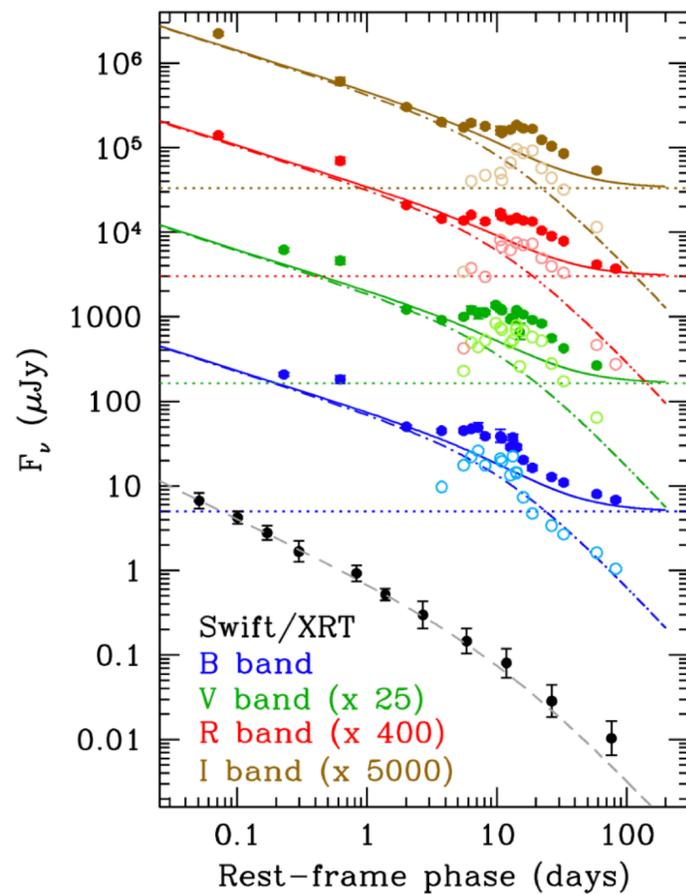
GRB 130702A/SN2013dx ($z = 0.145$)
D'Elia et al. 2015



TNG *Ugriz* (+5 spectra)+ REM *gr* obs.

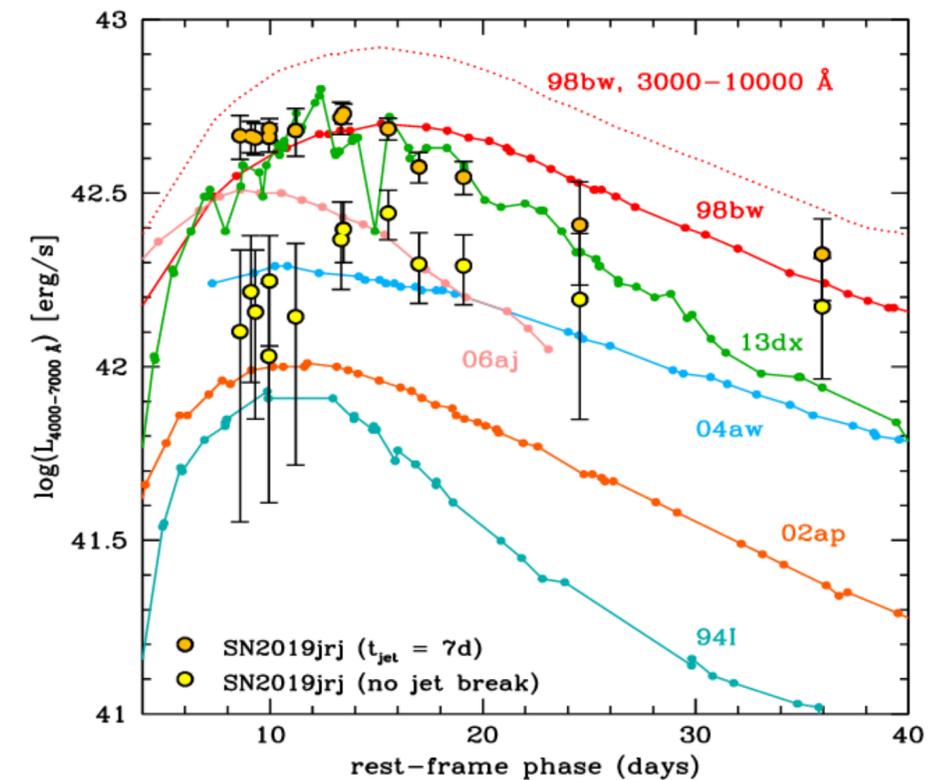
GRB-SN connection: ordinary events, ordinary monsters and real monsters

GRB 161219B/SN2016jca ($z = 0.148$)
Ashall et al. 2019



TNG *BVRI* obs.

GRB 190114C/SN2019jrj ($z = 0.4245$)
Melandri et al. 2021 submitted

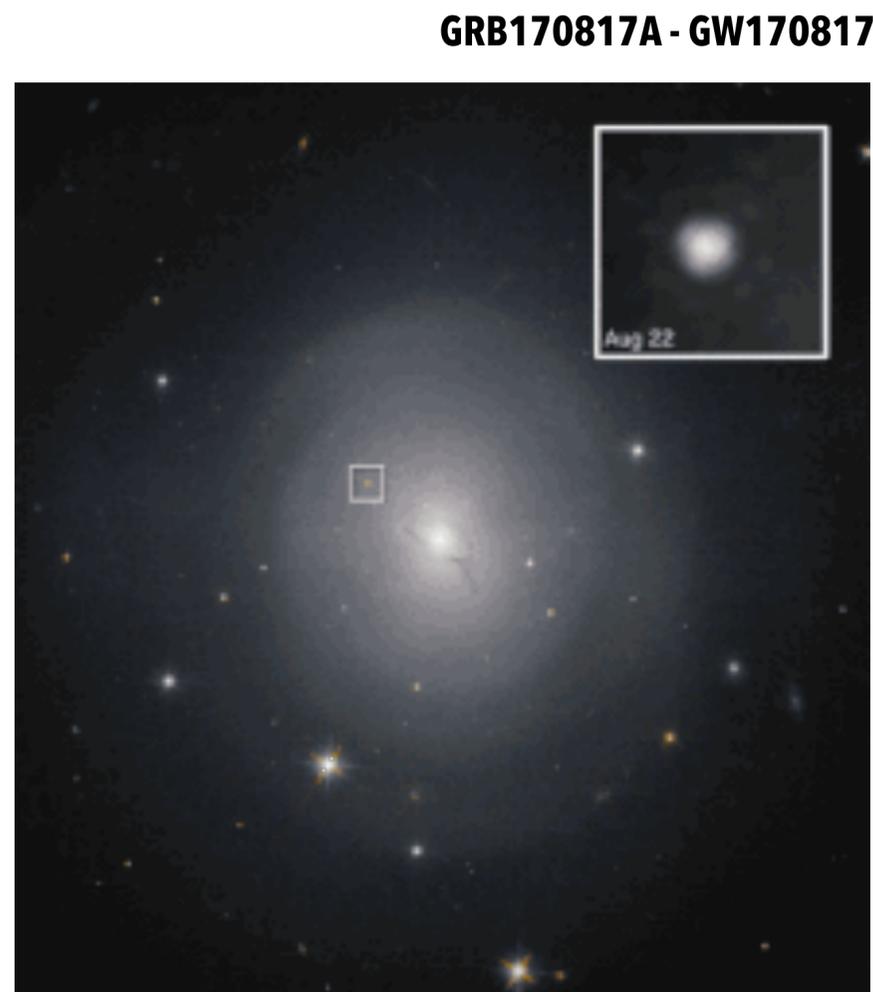


TNG *griz* + LBT *gri* obs. (+others)

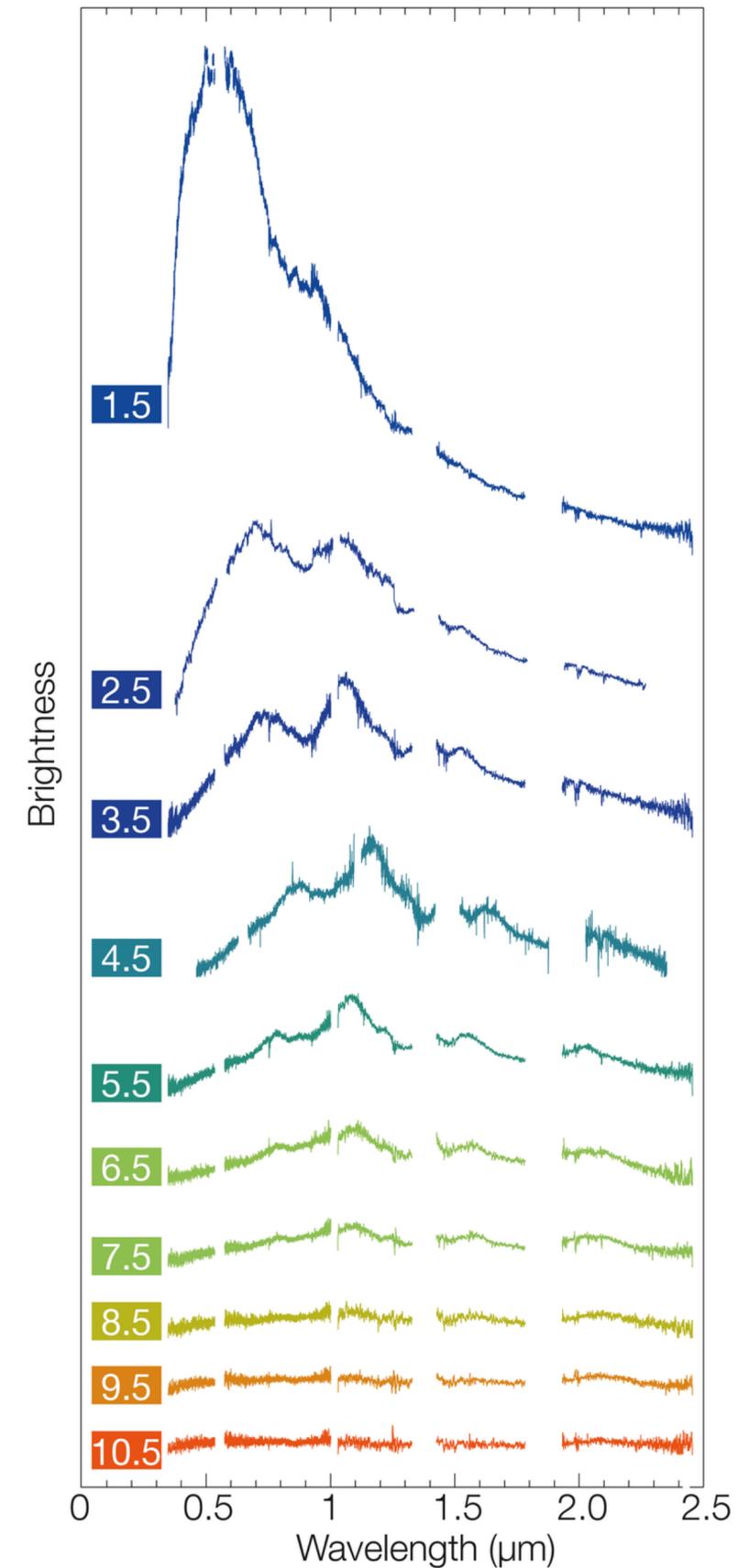
Short GRBs

Short GRBs are tricky to observe because fainter (even in the X-rays):

- z determination is critical
- different environment and energetics
- strong link with GW (progenitors)



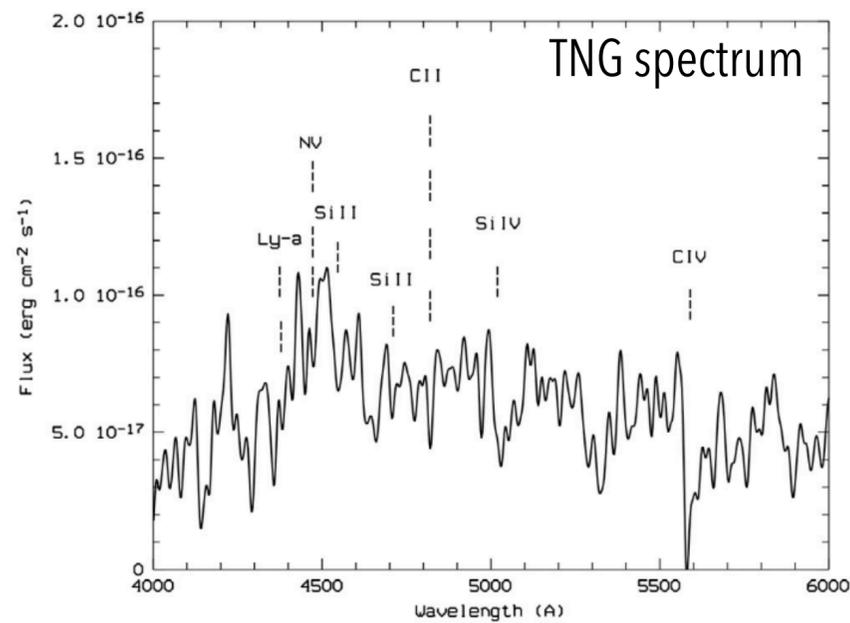
Credits: ESO



Credits: Pian, D'Avanzo +17, Smartt +17

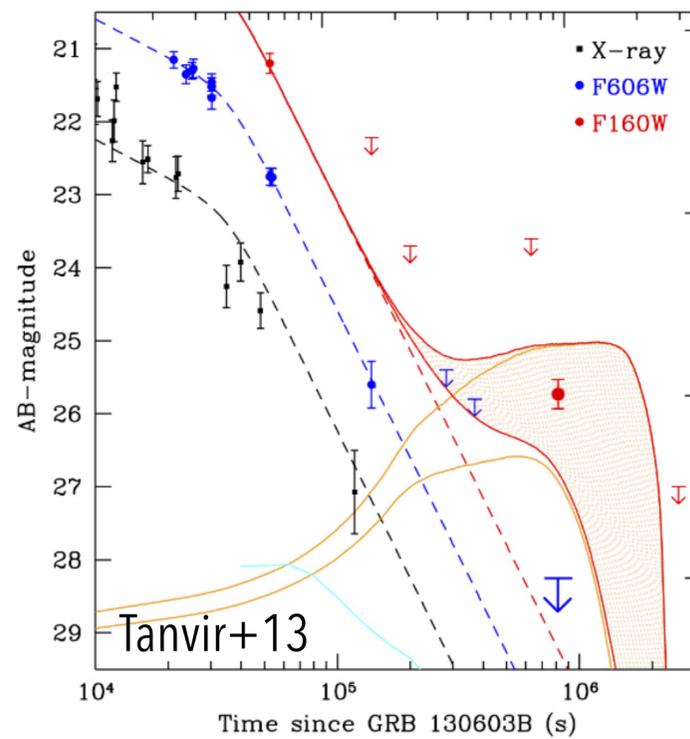
Short GRBs

GRB 090426 ($z = 2.609$)
Antonelli et al. 2009



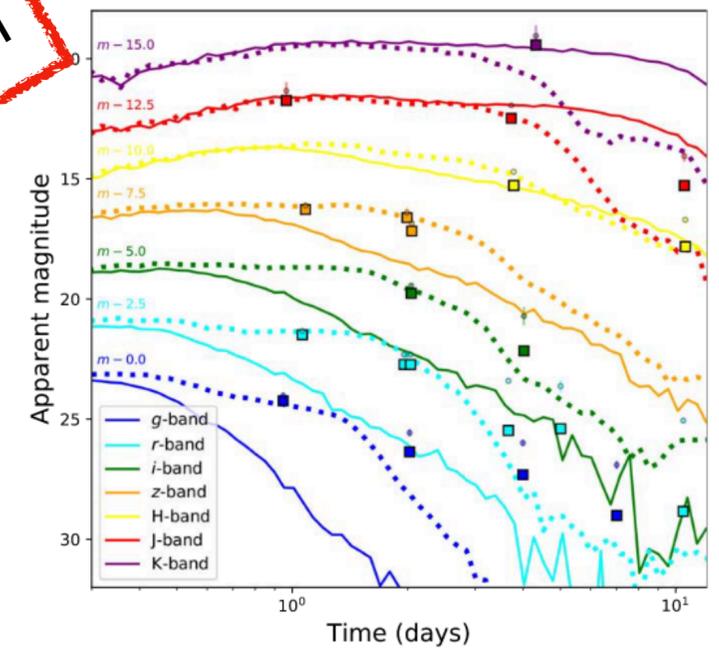
TNG *BVRI* + LBT *griz*

GRB 130606B ($z = 0.3565$)
de Ugarte Postigo et al. 2014



TNG *ri* obs.

GRB 160821B ($z = 0.162$)
Lamb et al. 2019

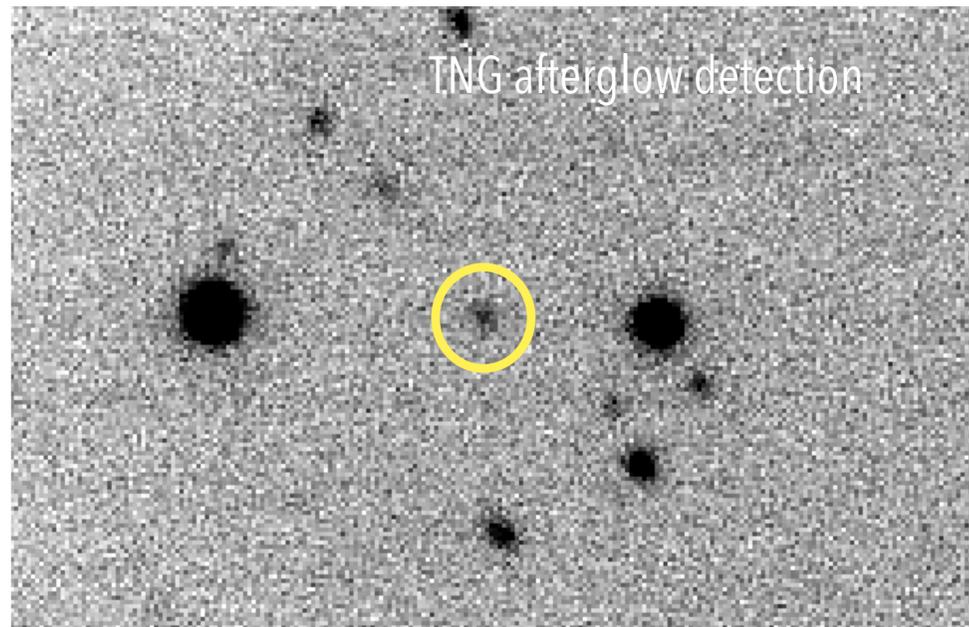


TNG *r* obs.

KN indication

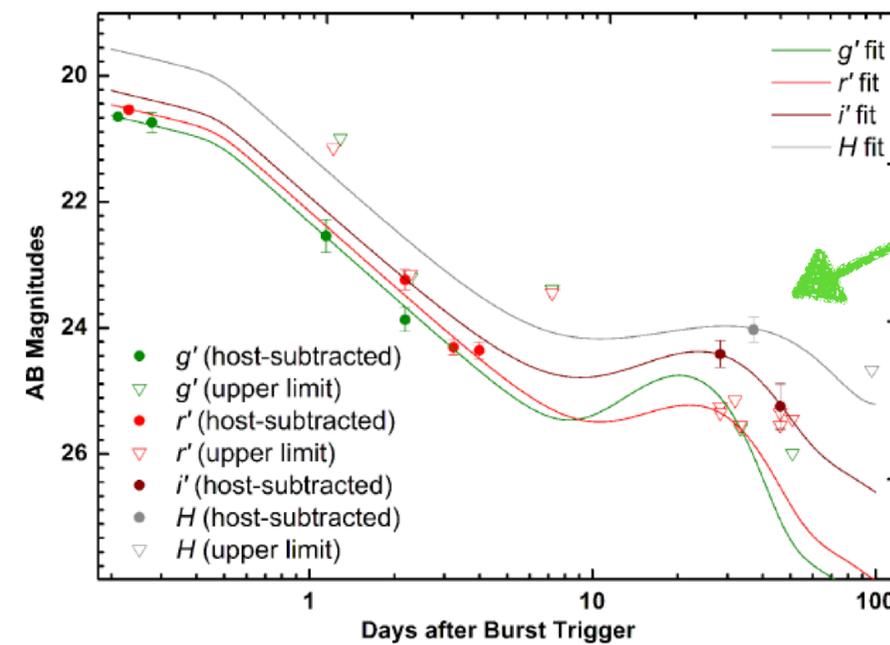
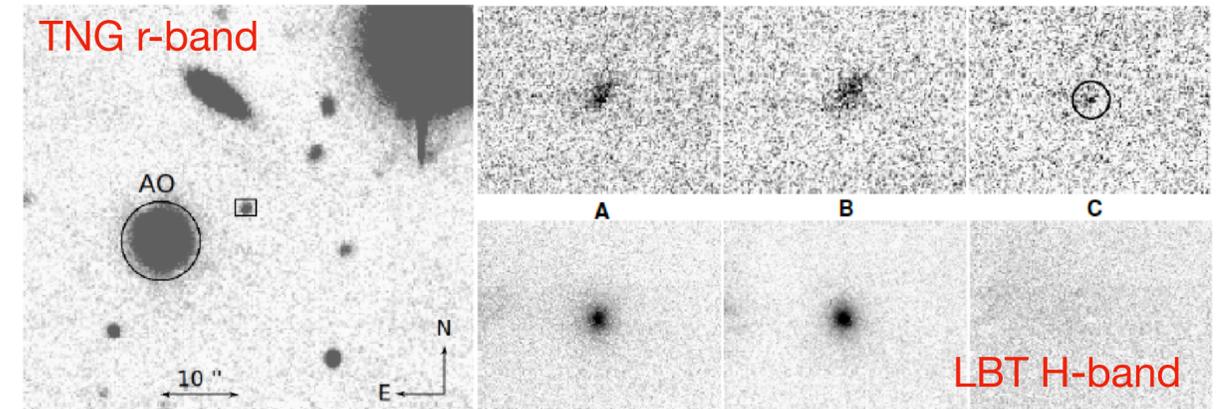
Short GRBs

GRB 160927A ($z = ??$)
GCN with TNG OA detection



TNG r obs.

GRB 200826A ($z = 0.748$)
Rossi et al. 2021



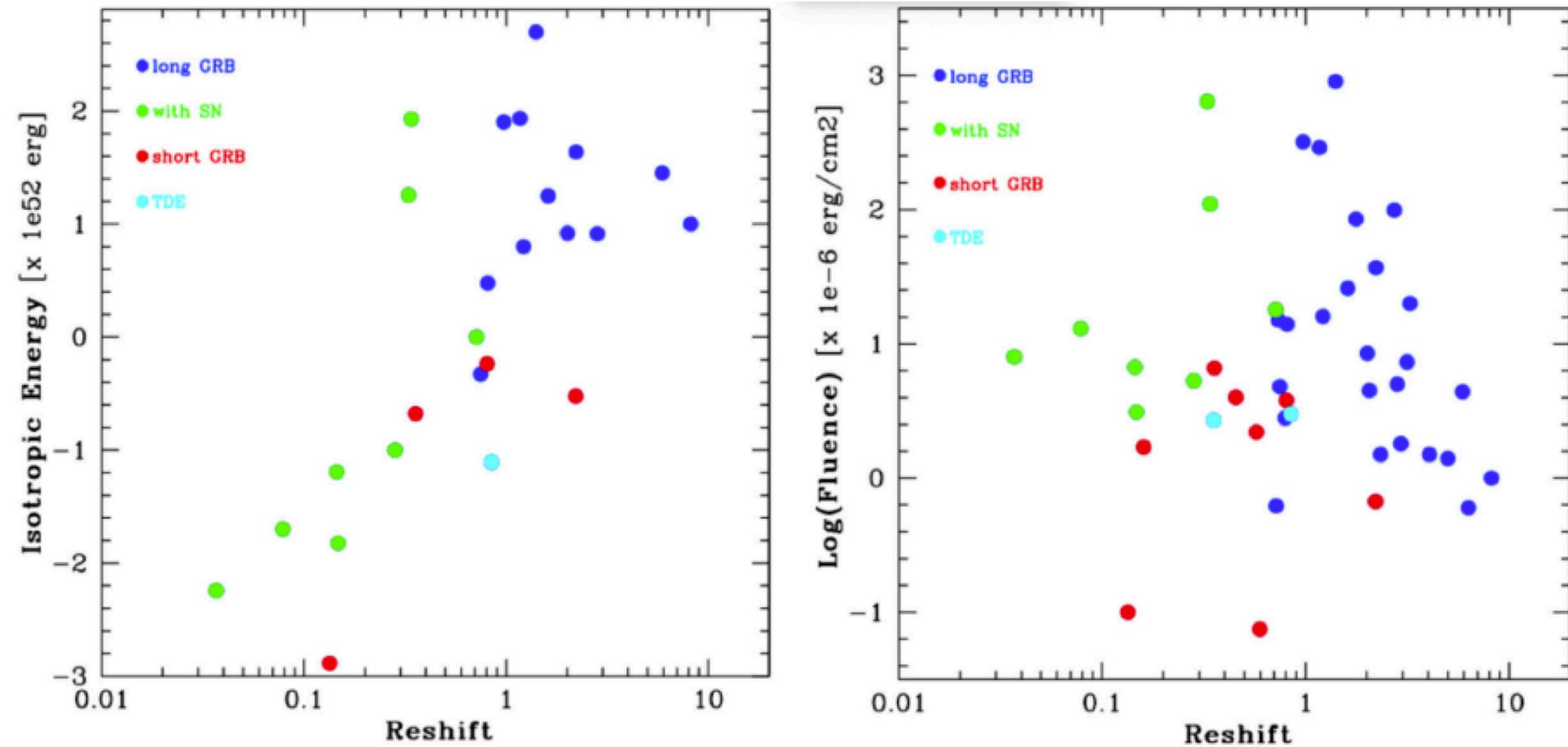
SN Bump?

TNG r +
LBT grH obs.

Conclusions: legacy value

Some numbers (since 2010)

- observed ~ 3 GRB/semester that met our four criteria (~ 65 overall)
- 62 GCN circulars
- 36 peer-reviewed papers, 1915 citations
(~ 3 papers/year with 52 citations/paper on average)
- 1 Nature, 1 Nat. Co.
- average time request of less than 30 hours/semester (now 17.5 hr)



Conclusions: the future of GRBs @ TNG

- GRBs have a high science impact in many astrophysical fields
- The use of TNG in GRB studies played (and is playing) a fundamental role in keeping the Italian community in a leading position (LTP approach and synergy with other facilities)
- **Unpredictable events**: for a promptly visible afterglow, geographic location can give an edge to TNG, also compared to larger telescopes, provided observations can be scheduled flexibly and rapidly
- **New science** in the time-domain astronomy era (*Swift*, *SVOM*, *SOXS*, *ZTF*, *BlackGEM*, *LSST*)
- Focus on **rare events**:
 - ▶ GW sources emit **short GRBs** (GW 170817A), hence their study is of paramount importance to guide the (much more difficult) observations of binary neutron star mergers
 - ▶ **high-z GRBs** (as already demonstrated with AMICI)
 - ▶ nearby GRBs with **associated SN**
- **High visibility and scientific return with relatively small time investment**